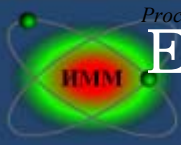


# Numerical simulation of turbulent stage of Richtmyer-Meshkov instability with multi-shock interaction.

*V.F.Tishkin, M.Ye.Ladonkina & N.V.Zmitrenko  
(IMM RAS)*

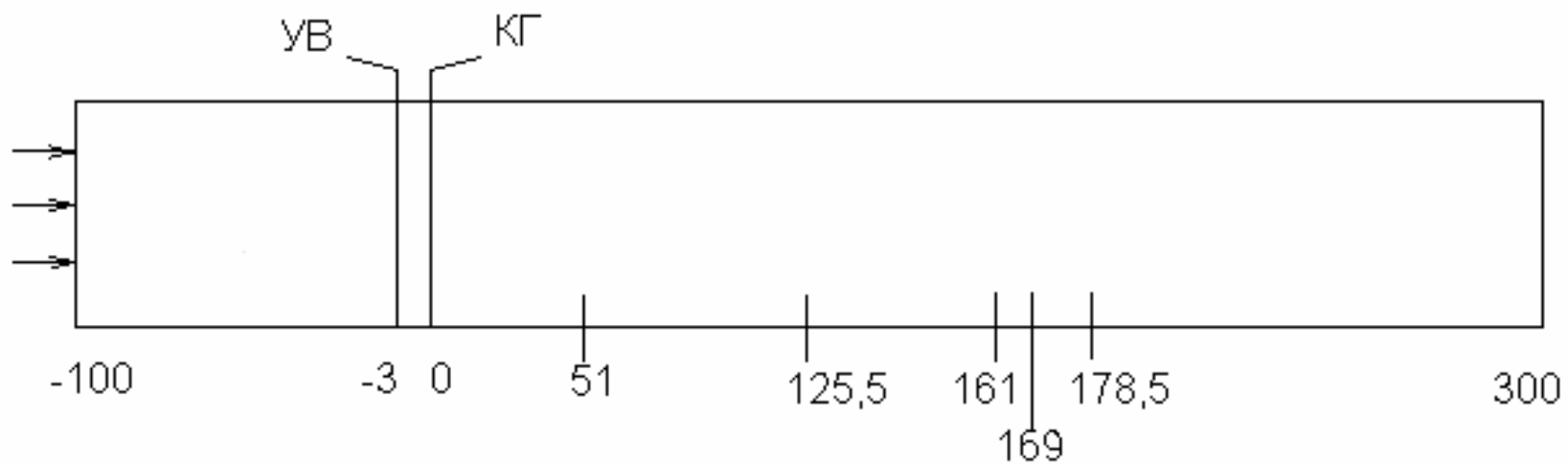
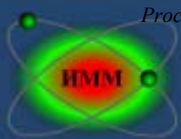
**9<sup>th</sup> International Workshop on the  
Physics of Compressible Turbulent Mixing**  
Cambridge, UK  
19-23 July 2004



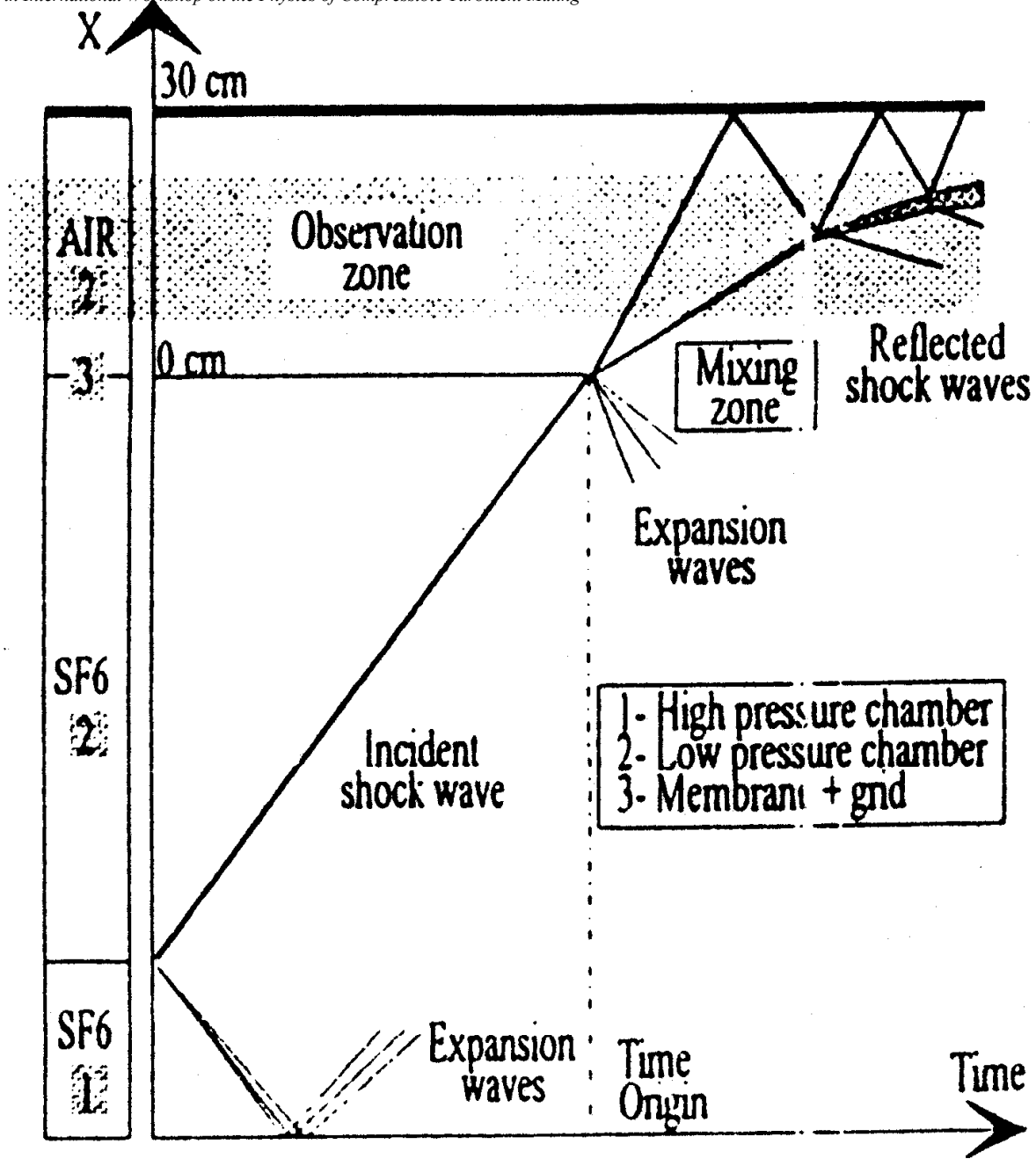
# Experimental [1] and numerical investigations of turbulent mixing

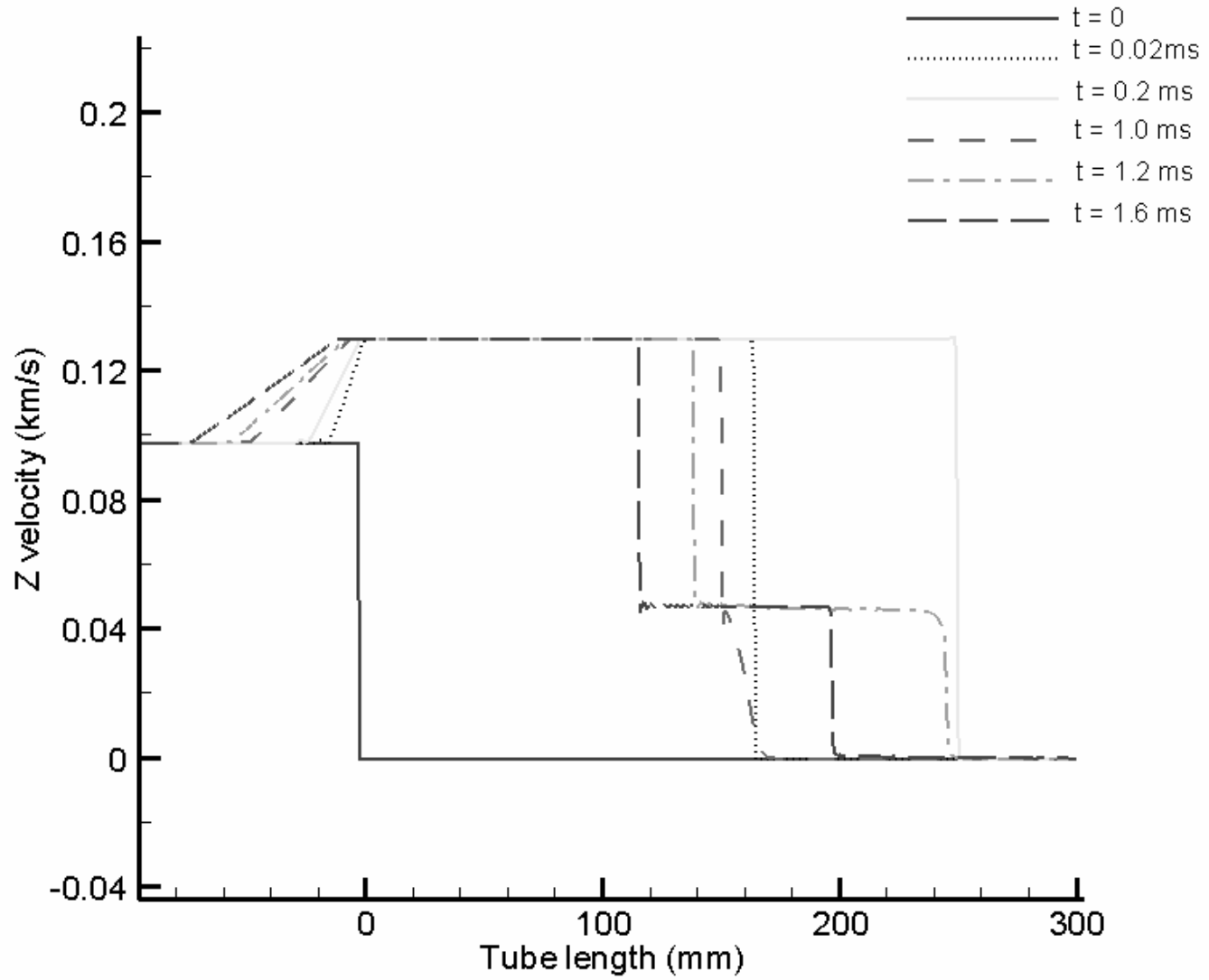
[1] F. Poggi, M.-H. Thorembeay, G. Rodriguez. Velocity measurements in turbulent gaseous mixtures induced by Richtmyer-Meshkov instability. // *Physics of Fluids*, 1998, Vol. 10, No. 11, pp. 2698-2700

Richtmyer-Meshkov instability development for SF<sub>6</sub>/Air (adiabatic exponents are 1,094 and 1,4)

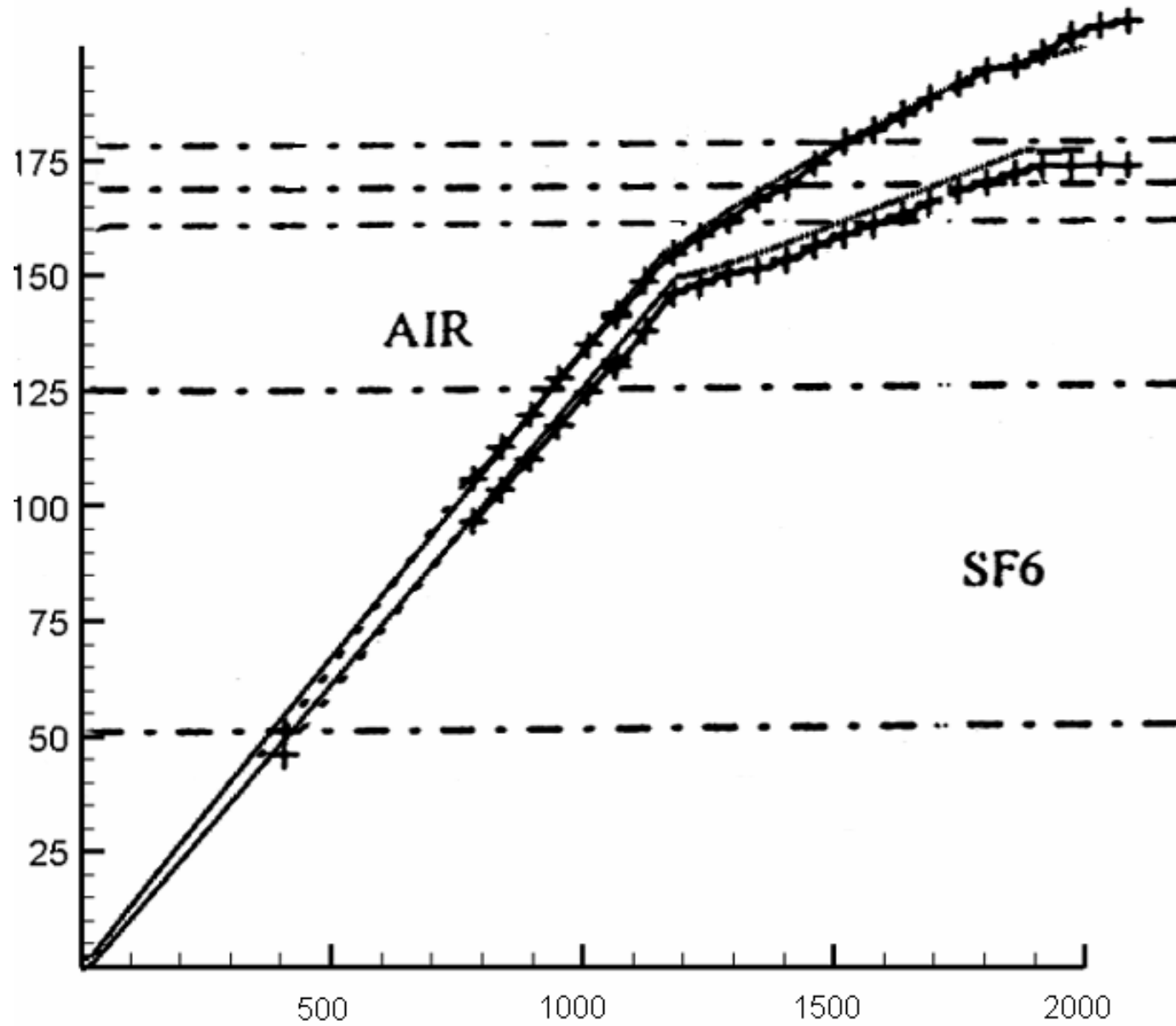


- Initial temperature 291<sup>o</sup>K
- Pressure in low-pressure camera 1Bar.
- Pressure after shock in SF<sub>6</sub> 2,152 Bar.
- Density 1,209·10<sup>-2</sup> g/cm<sup>3</sup> ,
- Shock speed 195,2 m/s, flow speed after shock 97,76 m/s,
- Initial density SF<sub>6</sub> in low-pressure camera 6,037·10<sup>-3</sup>g/cm<sup>3</sup>,  
air - 1,198·10<sup>-3</sup> g/cm<sup>3</sup>
- Shock speed in air 428,7 m/s
- Contact boundary speed 130,1 m/s.
- SF<sub>6</sub> and air are ideal gases

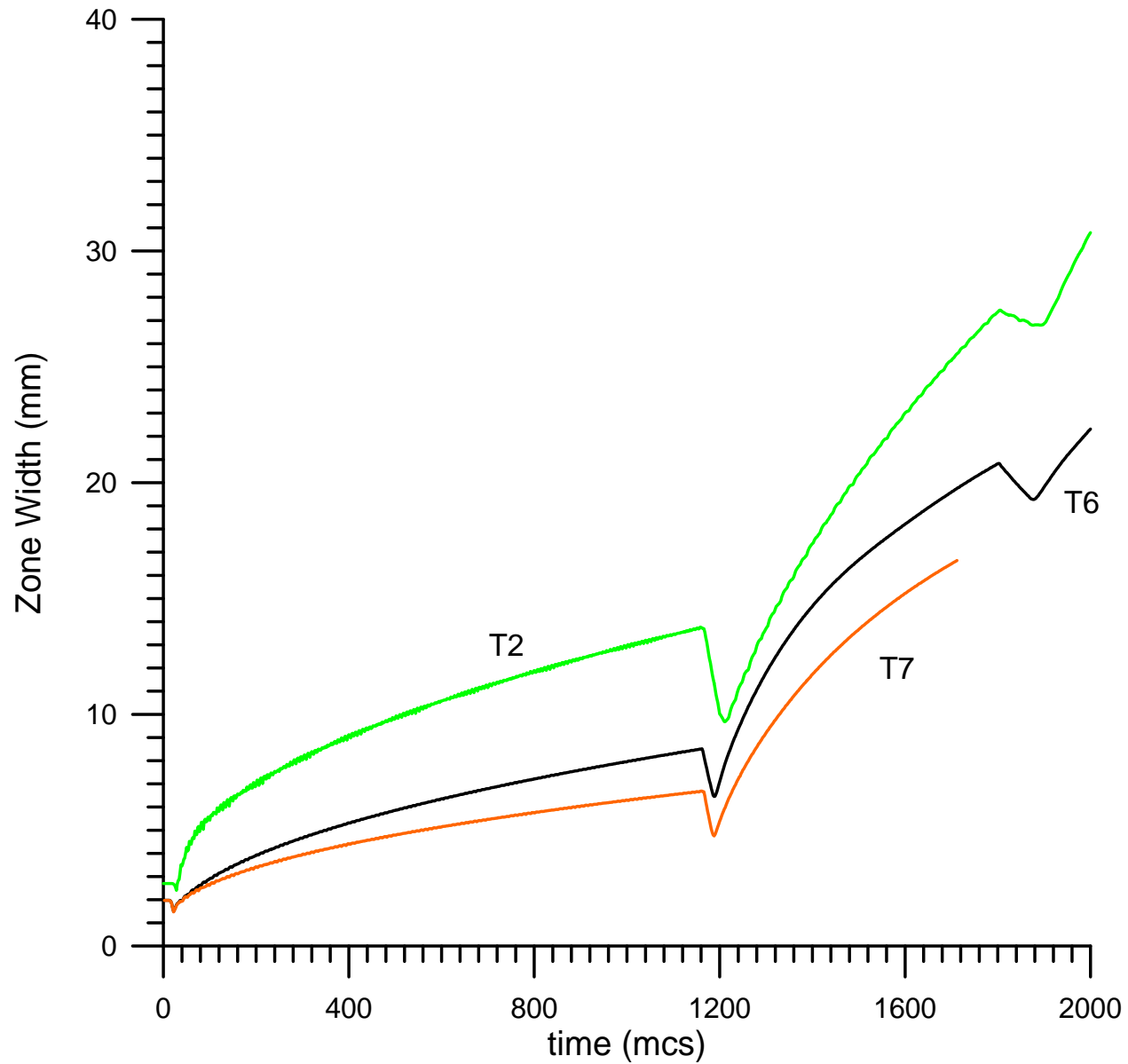


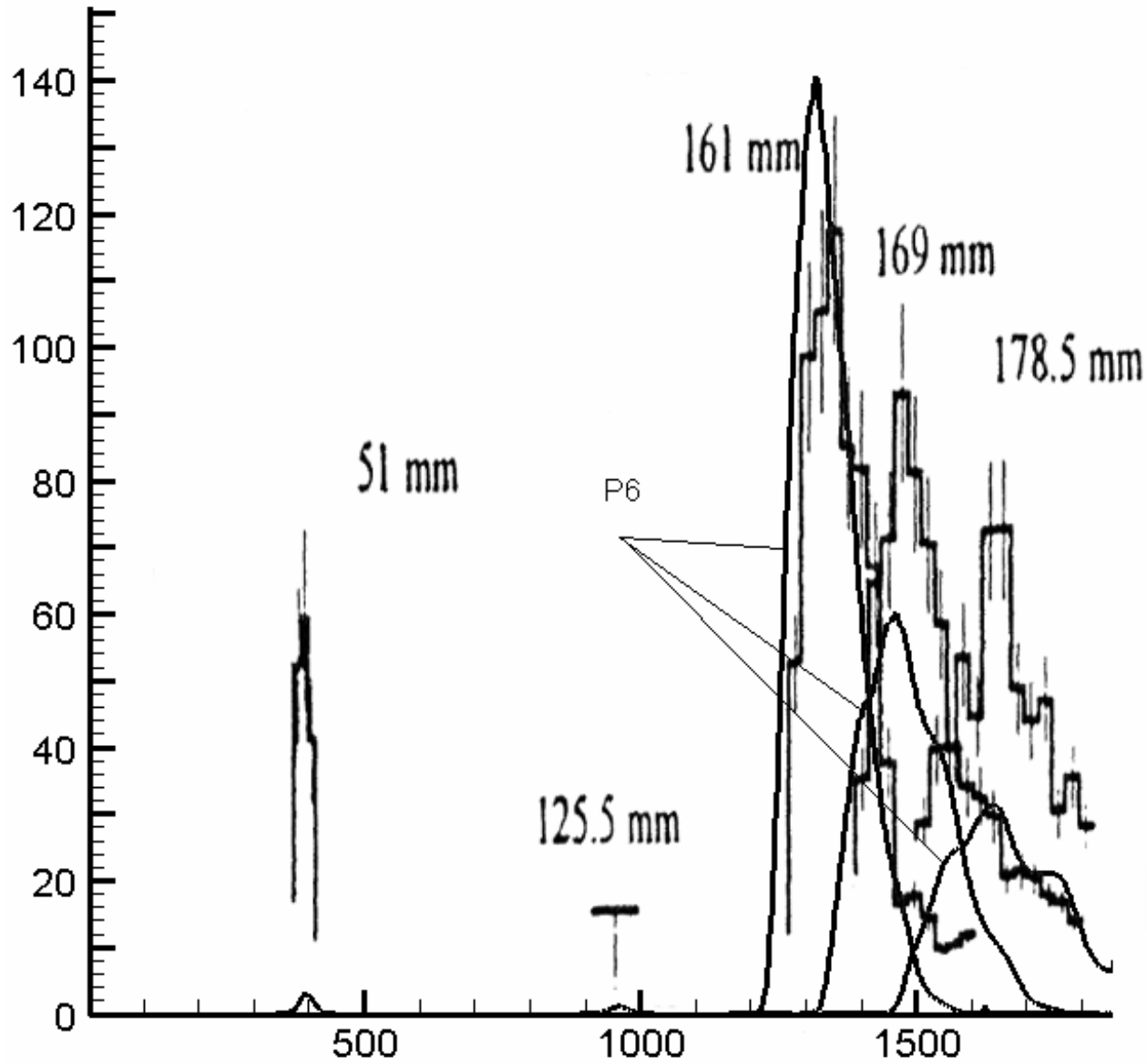


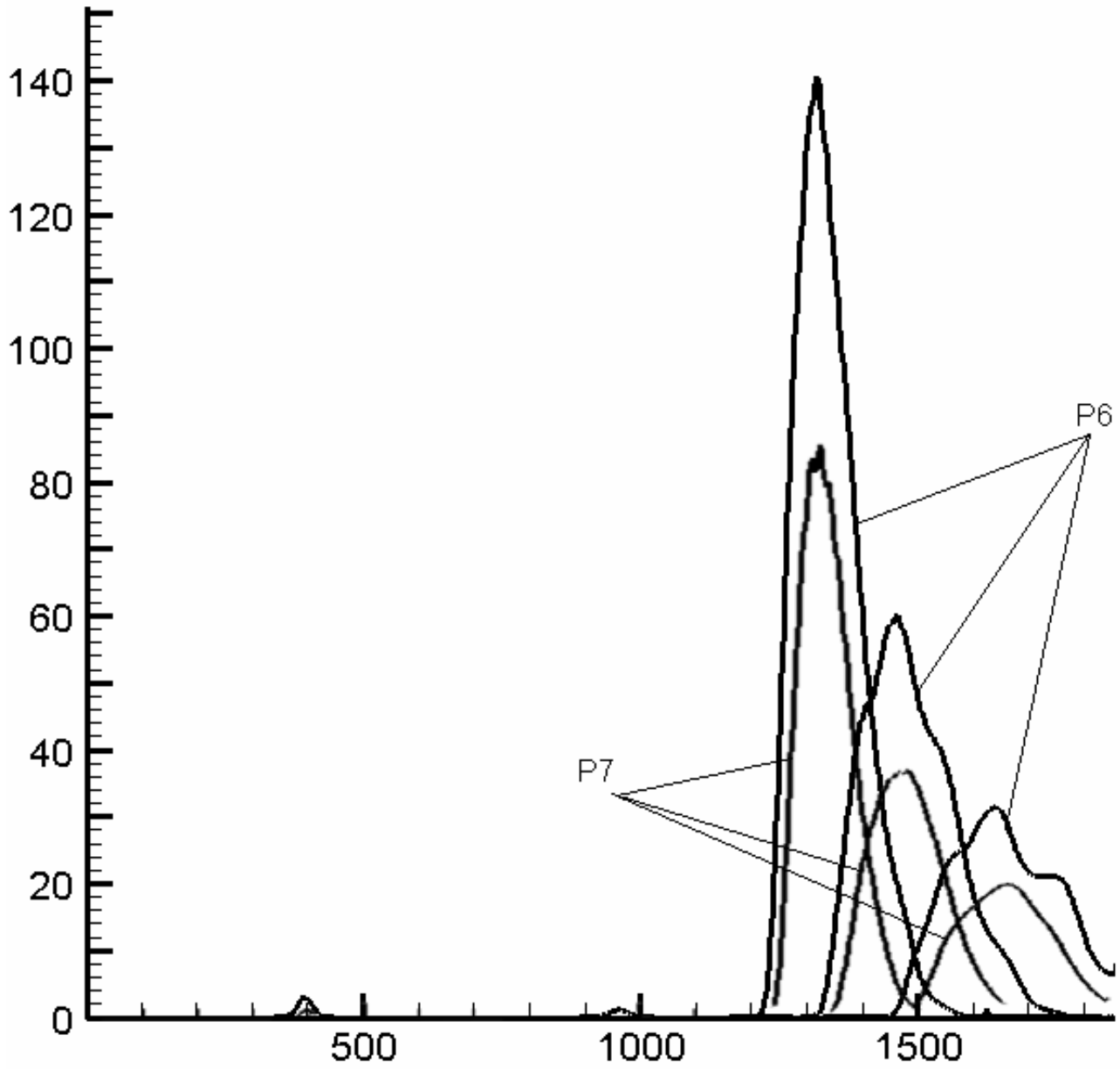
- On the basis of the calculations executed on grids with  $h=1$  and  $0,5\text{mm}$  it is possible to draw the following conclusions:
- For satisfactory reproduction of experimental data [1] it is enough to carry out calculations in area with the sizes  $0 \leq x \leq 80$  mm,  $0 \leq y \leq 80$  mm, -  $100 \leq z \leq 300$  mm; Before arrival of a shock wave on contact border in a zone of the unit of gases there are no perturbations of numerical character.
- The sizes of cross-section section of a tube do not render appreciable influence on results of calculations.
- The type of the task and size of initial perturbations of density noticeably affect results of calculations.
- For the further calculations we shall reduce a step of a grid ( $h=0.333$  mm) and we shall keep a level of initial perturbations in a transitive layer (20 % of a variation of density with the additional requirement  $0 \leq C \leq 1$ ).
- These calculations can be lead for section of a tube  $4 \times 4$  sm that will allow to reduce the general number of accounting cells twice.

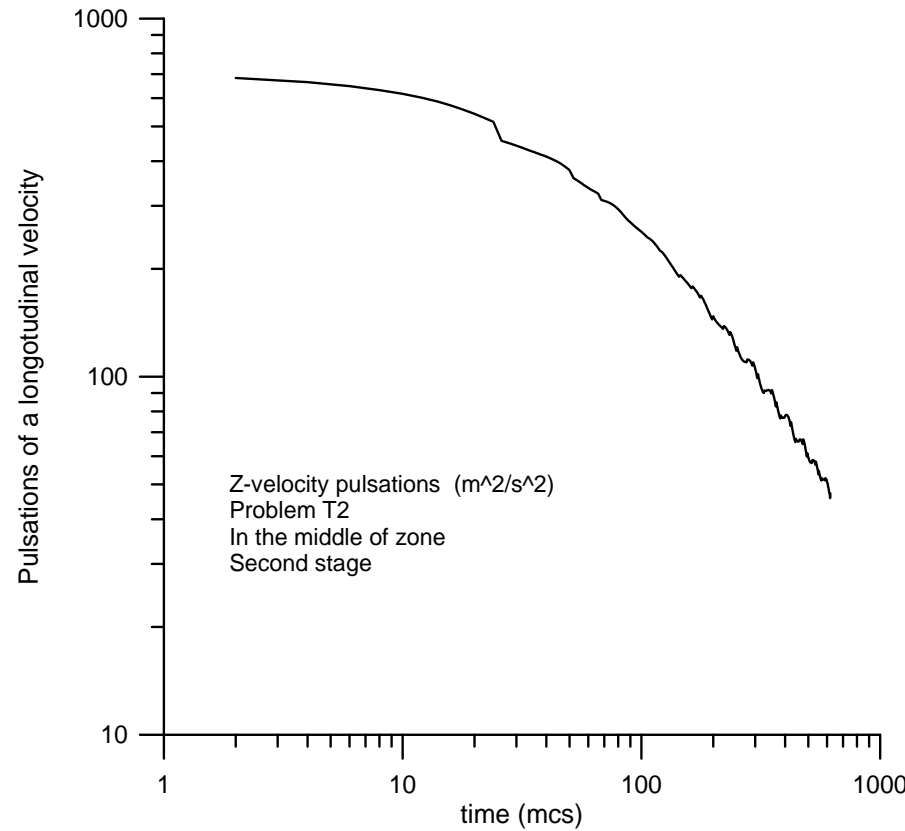
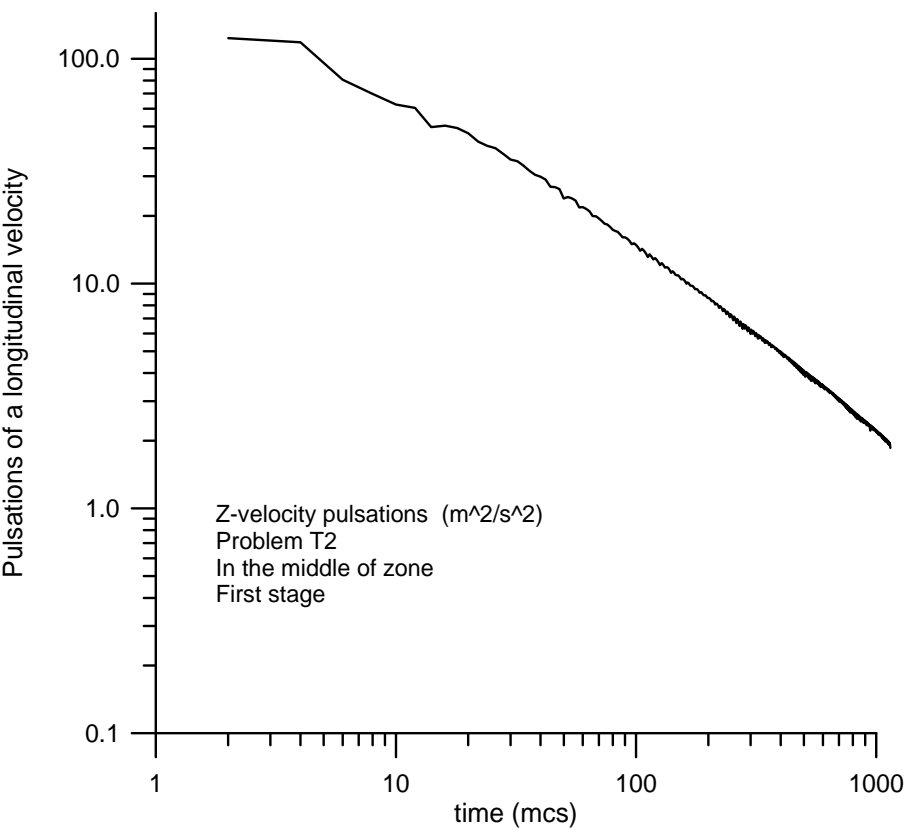
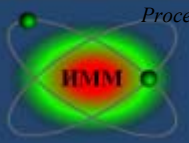


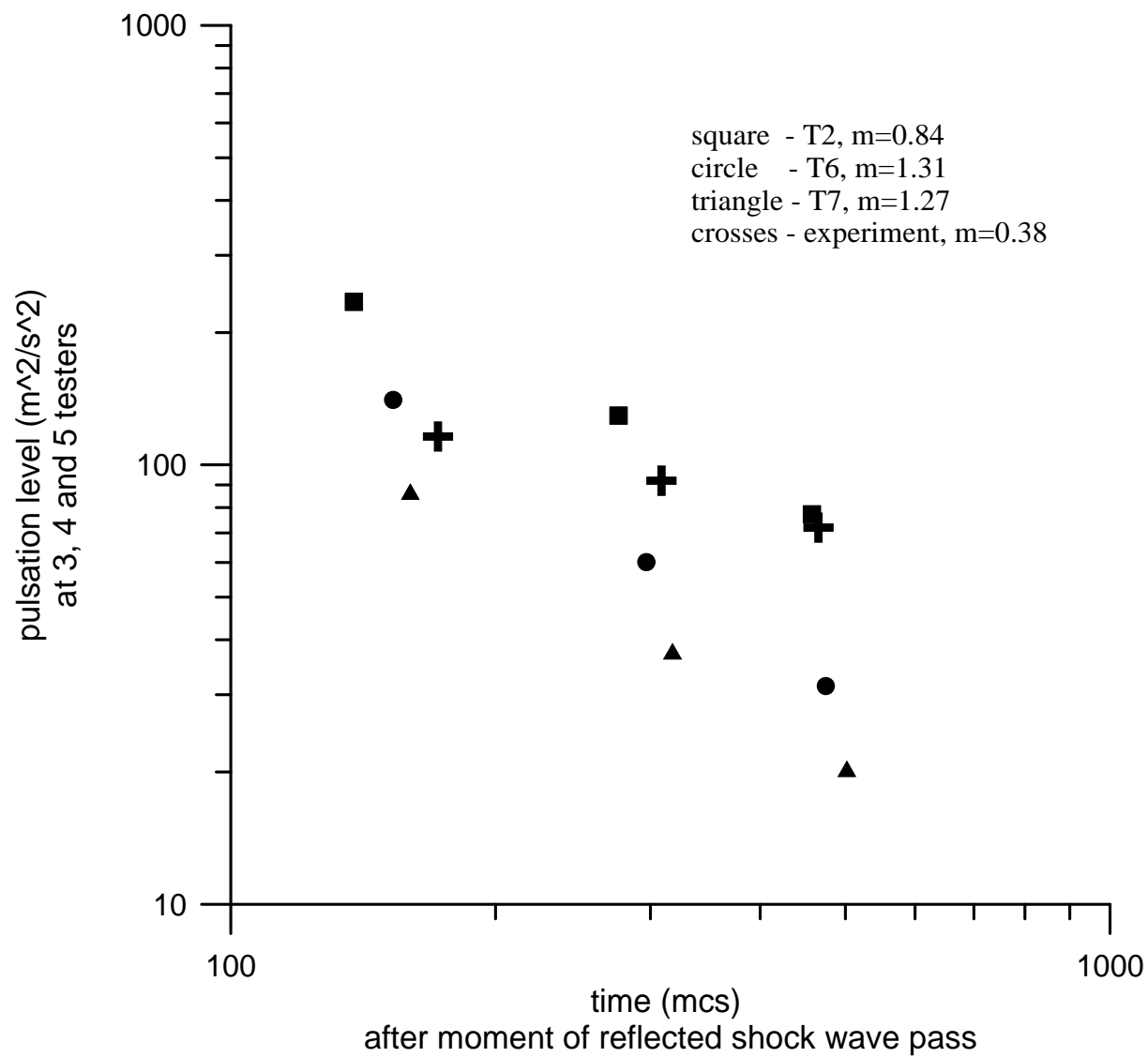
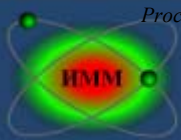










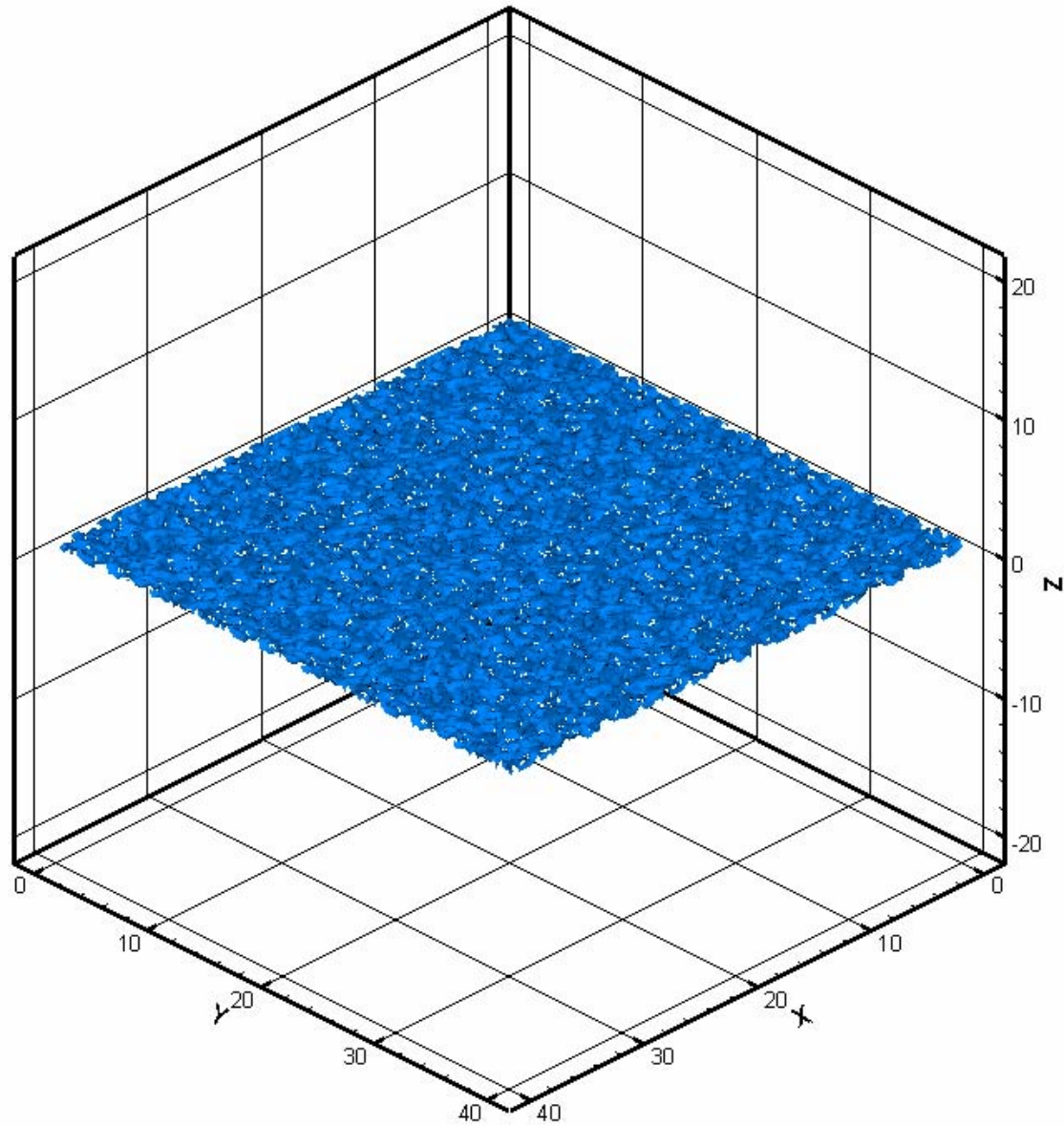


- Pulsation in mixing zone  $W_p(t)$  may be approximate by expression

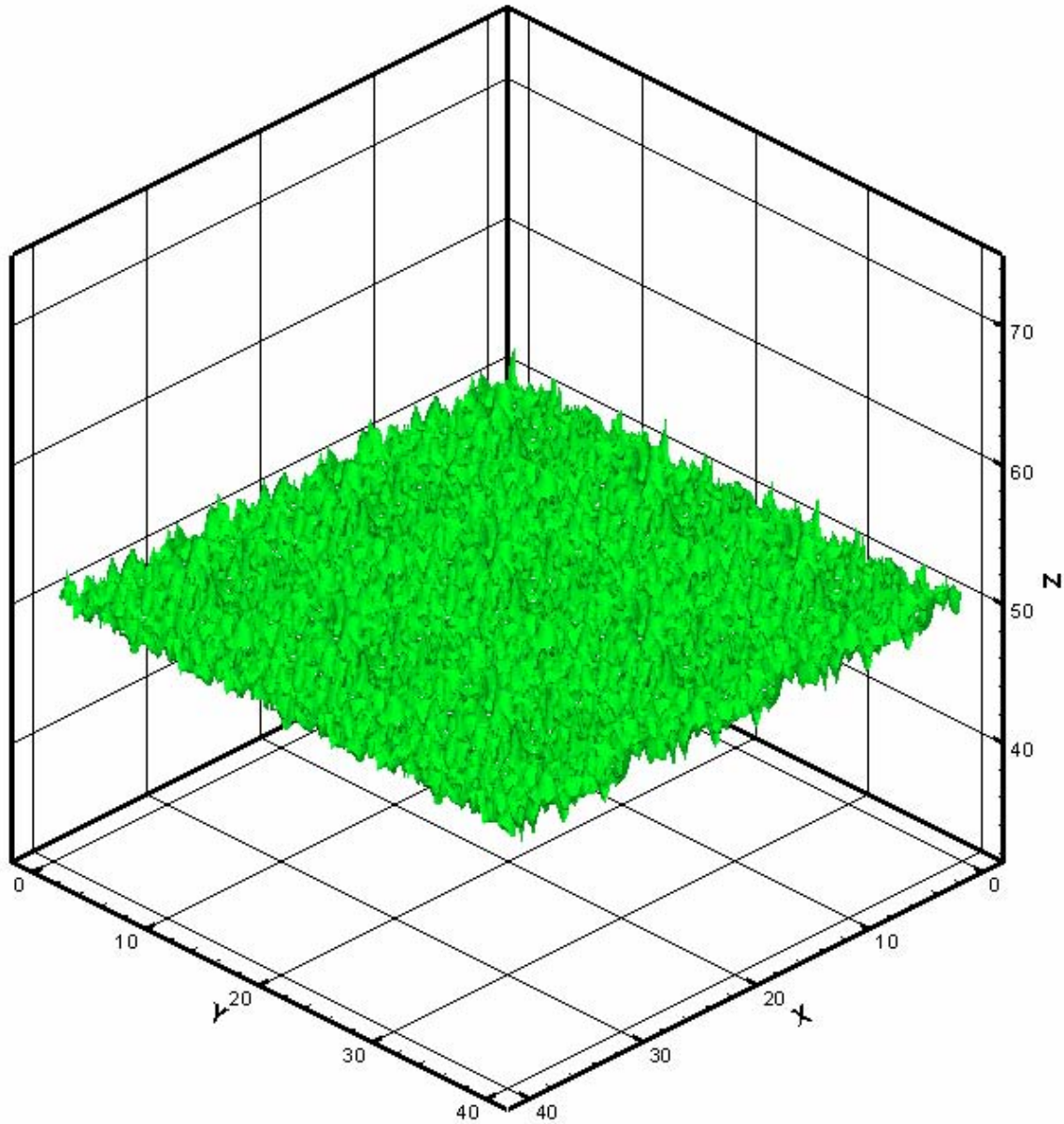
$$W_p(t) = B(t-t^*)^{-m},$$

$$W_p \sim (dL/dt)^2 \sim (L/t)^2 \sim (t-t^*)^{(2k-2)}, \quad m = 2 - 2k.$$

$t = 0.$

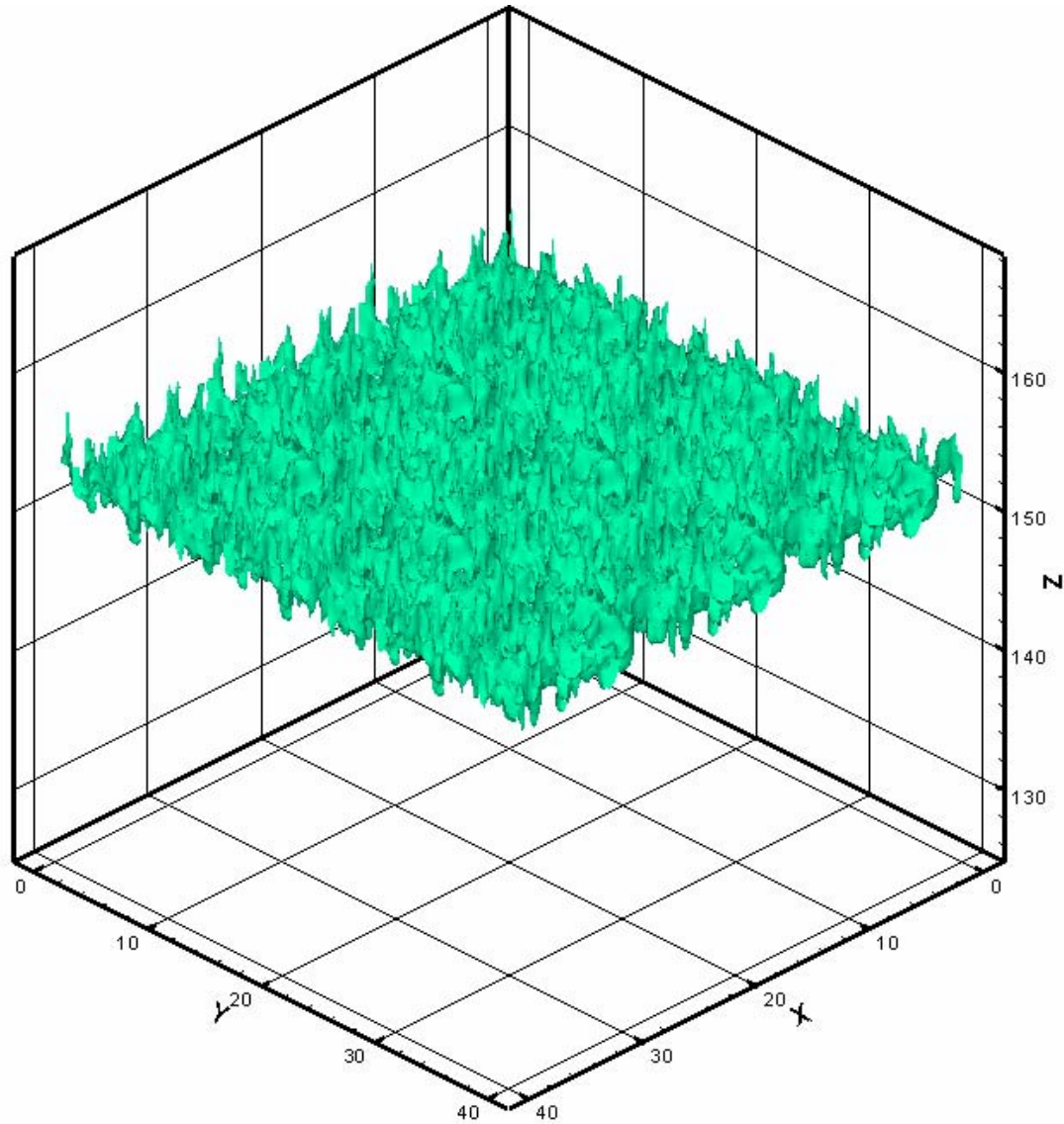


$t = 400 \text{ mks}$

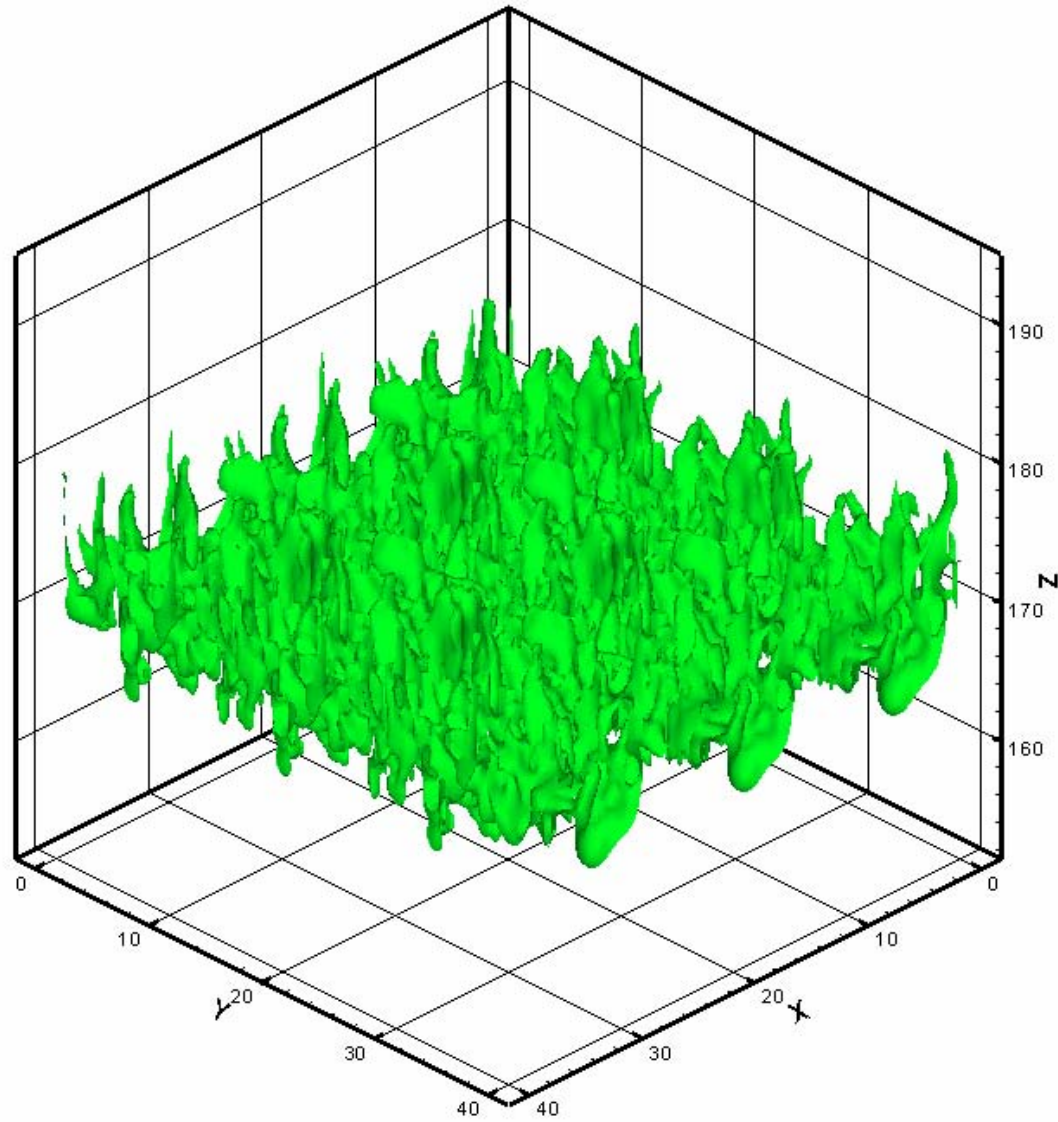




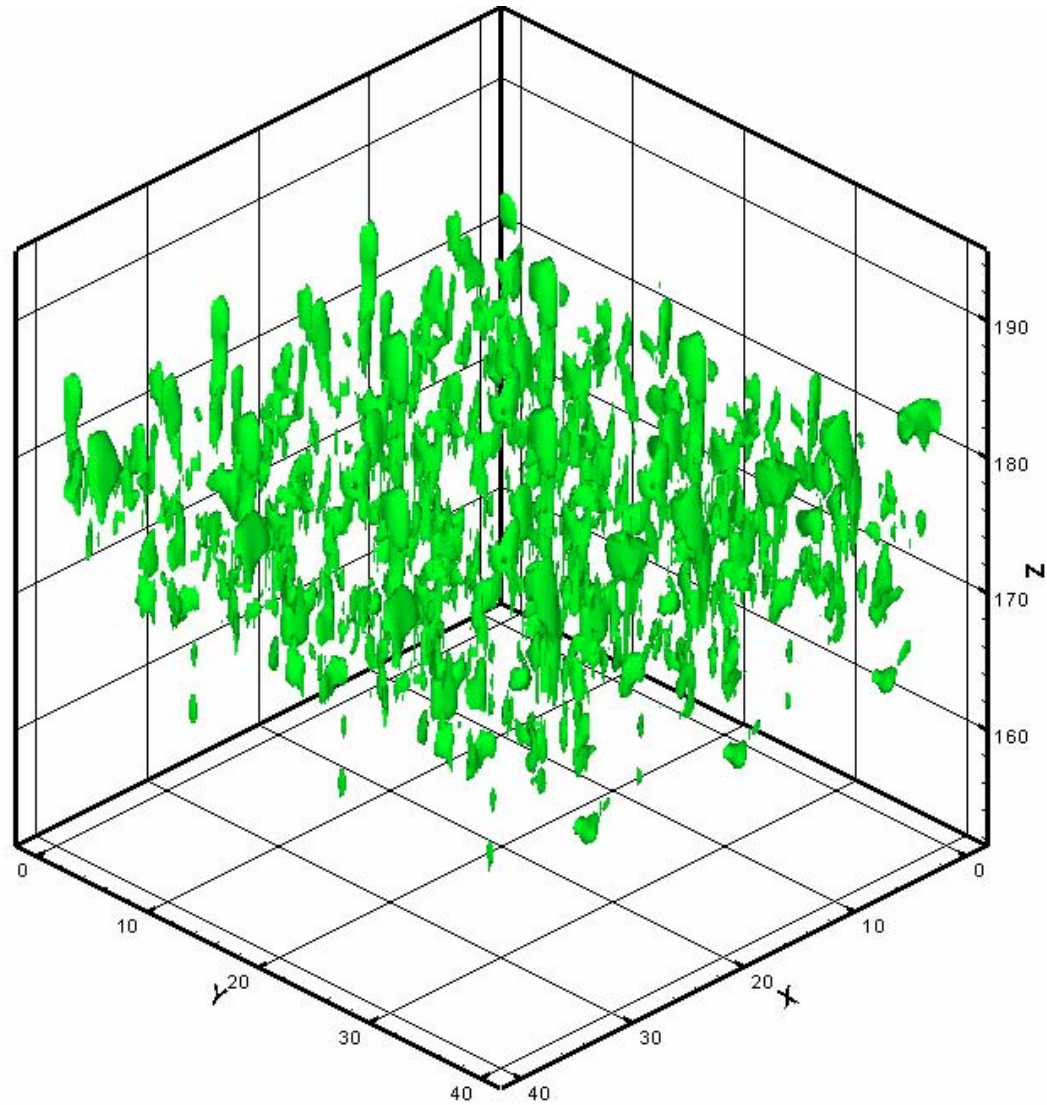
$t = 1200$  mks



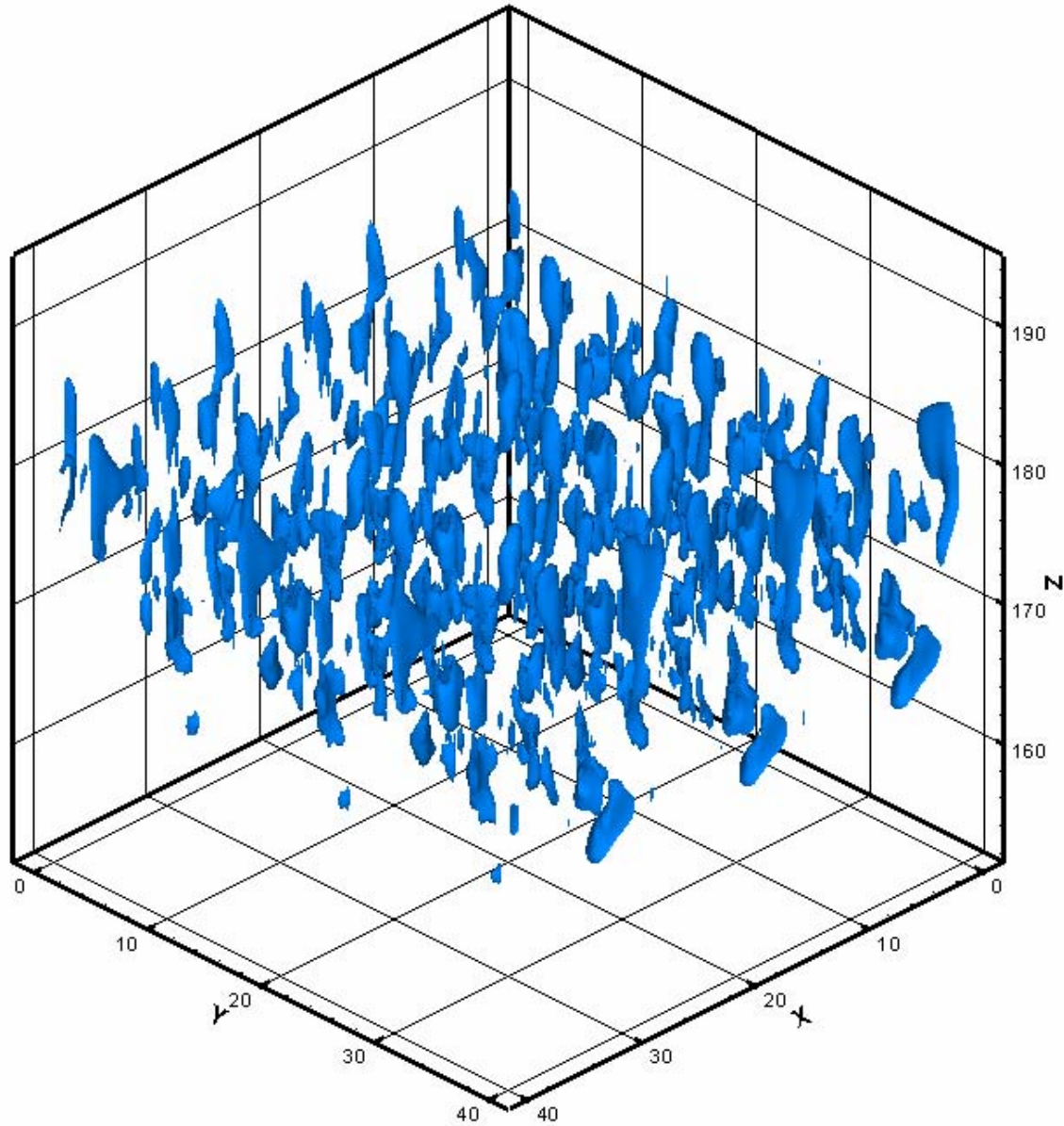
$t = 1600\text{mks}$



# Iso-Surface $|\text{rot}V|$ , $t=1600$ mks

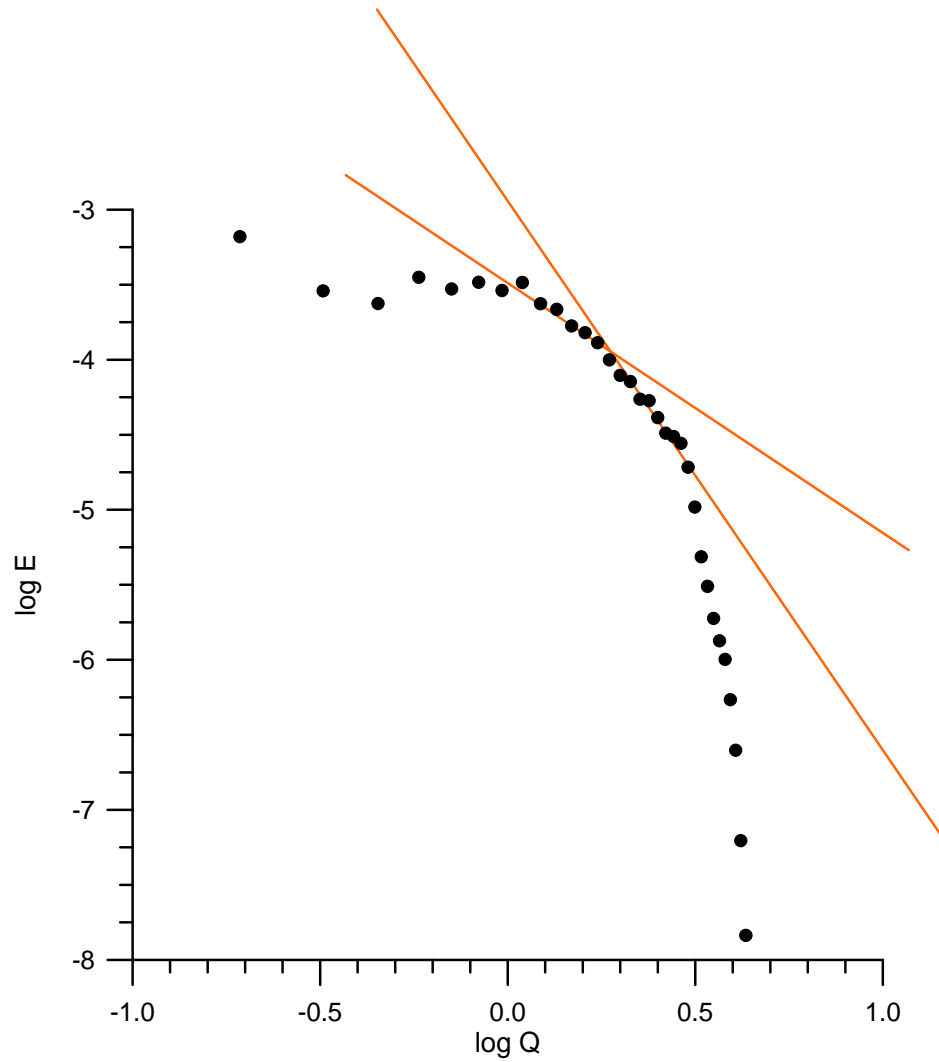


# Iso-Surface Ekin, t=1600 mks

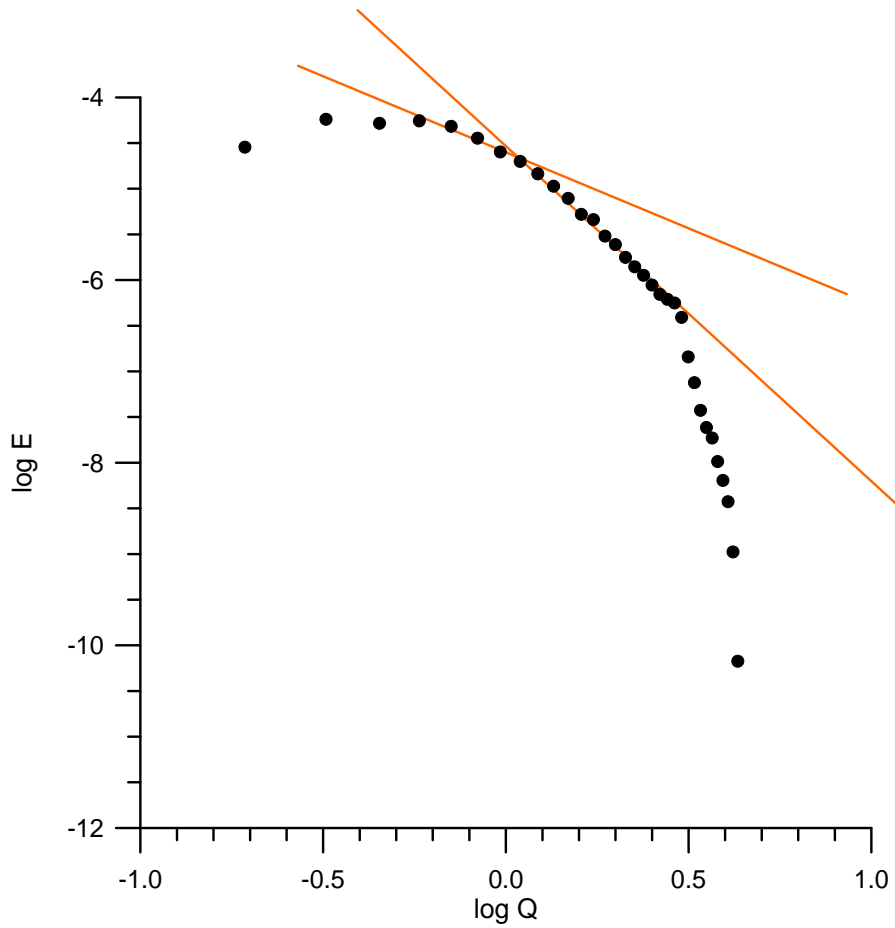


$t=1200mks$

T6



t=1600mks **T6**



**T7**

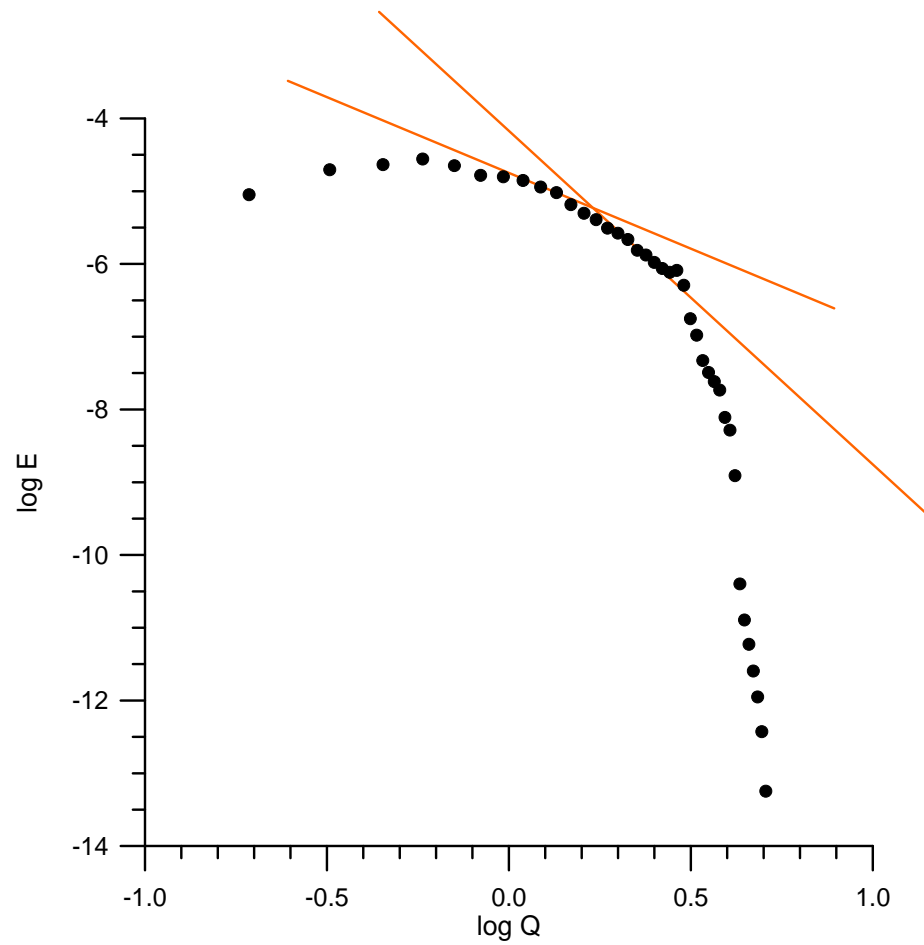
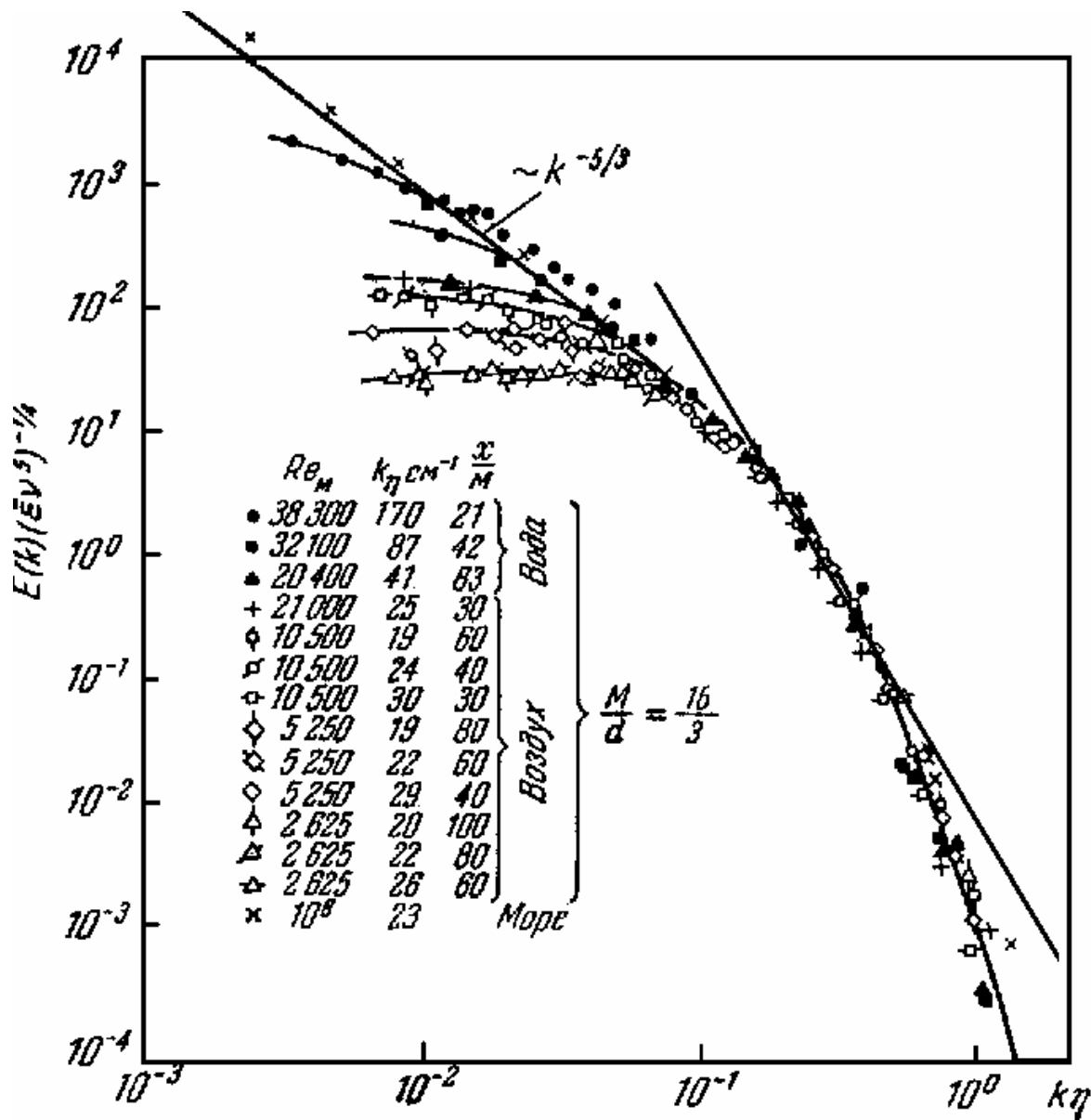
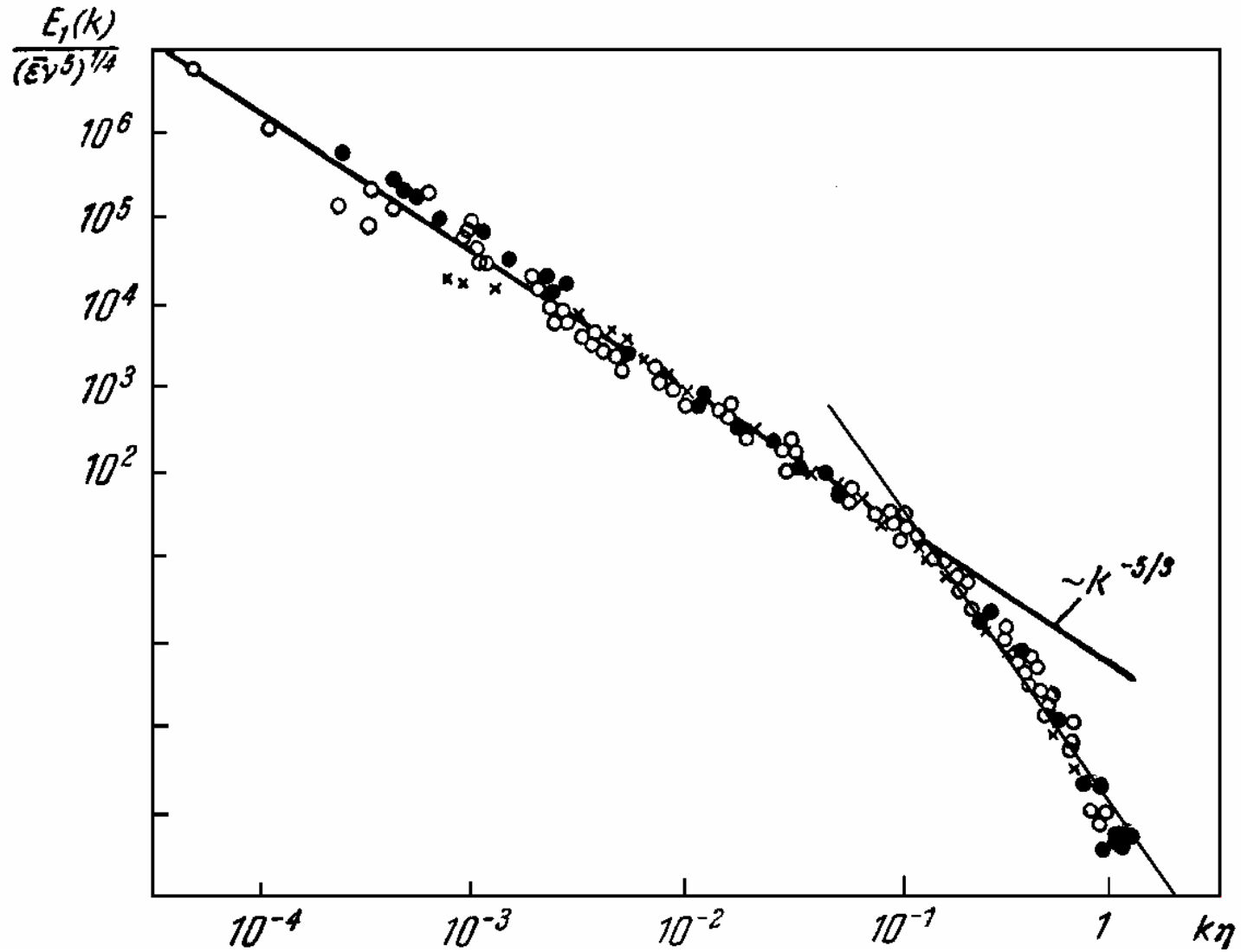


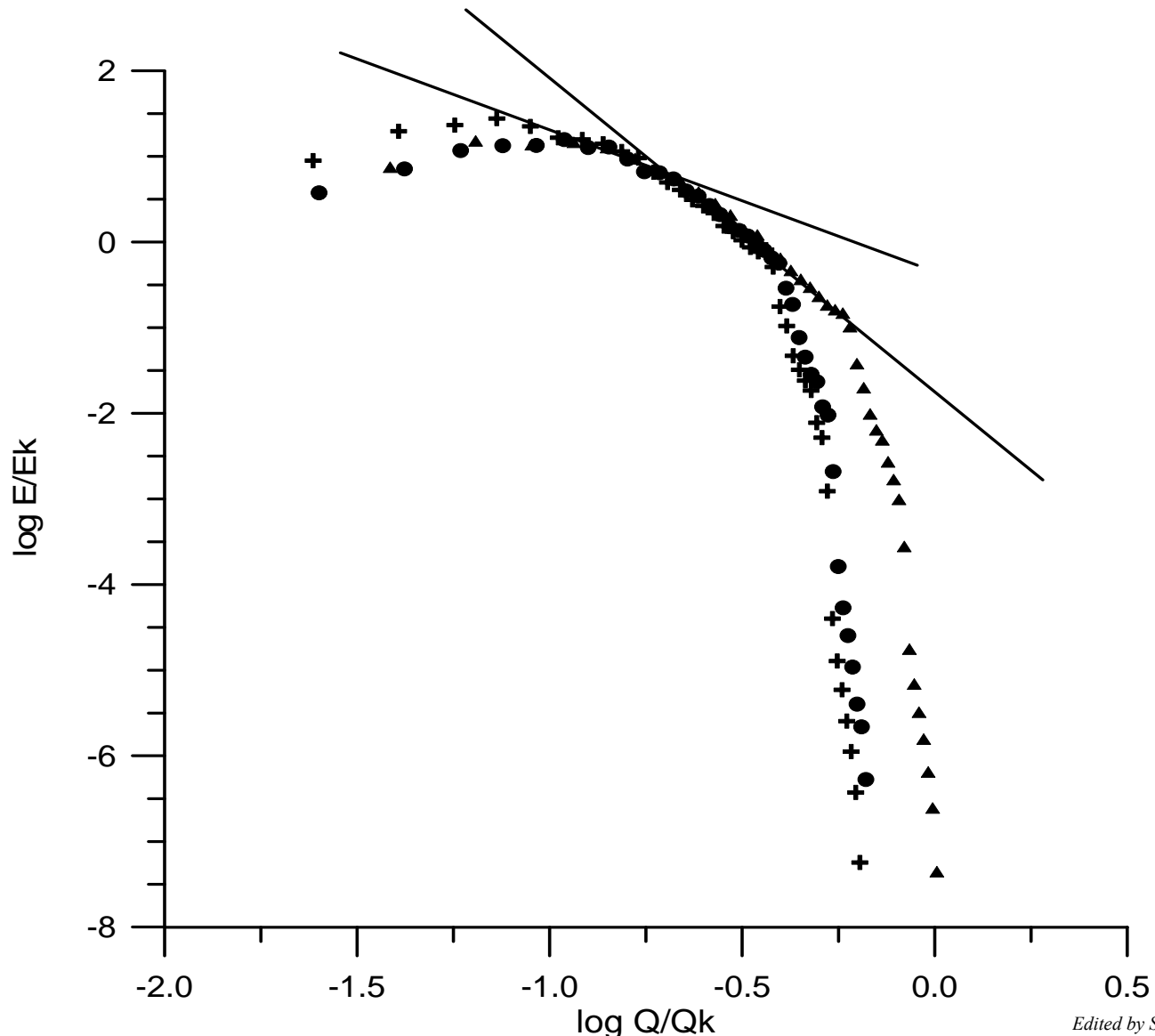
Figure from A.S.Monin and A.M.Jaglom monograph. Experimental value of spectral density. Line have inclination  $-5/3$  and  $-11/3$ .







Orthonormalized spectral density of kinetic energy in tasks T2 (circles), T6 (triangles) и T7 (crosses). Line have inclination  $-5/3$  and  $-11/3$ .

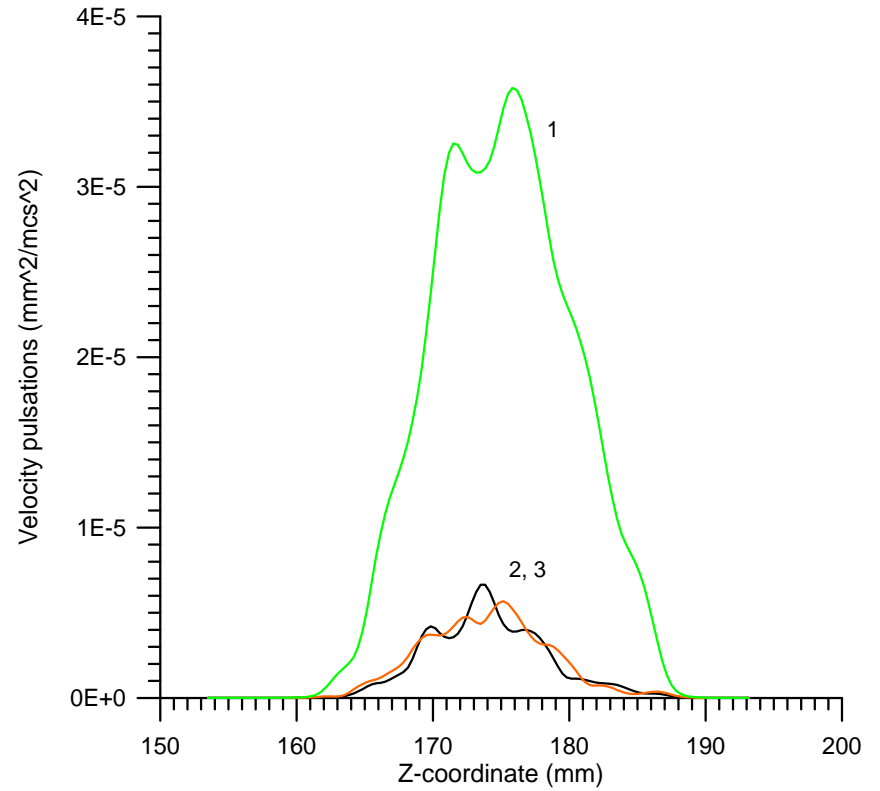
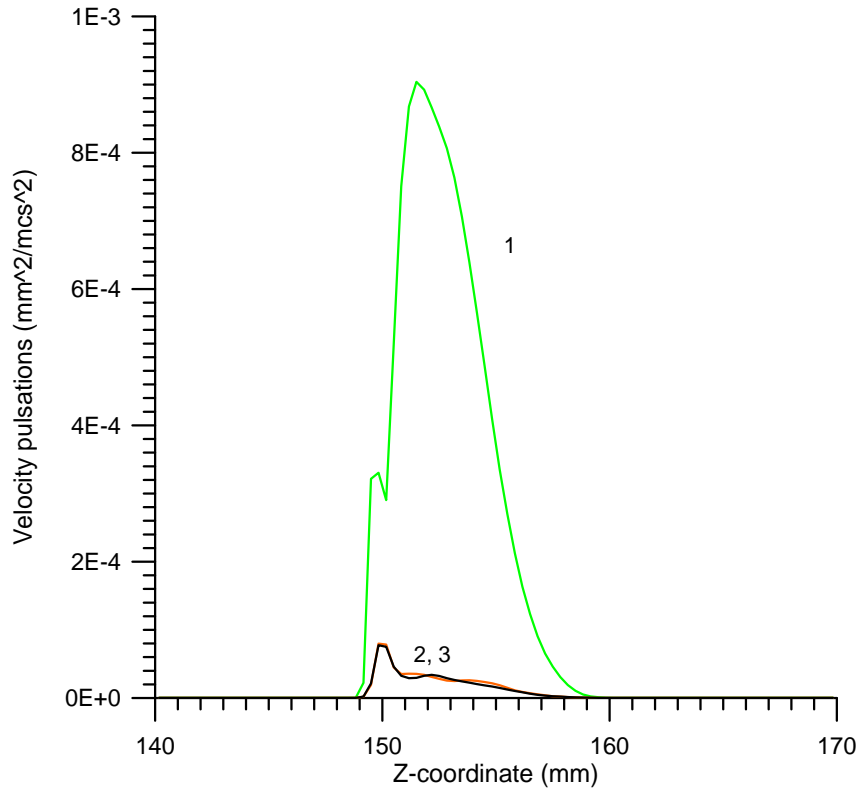


T6

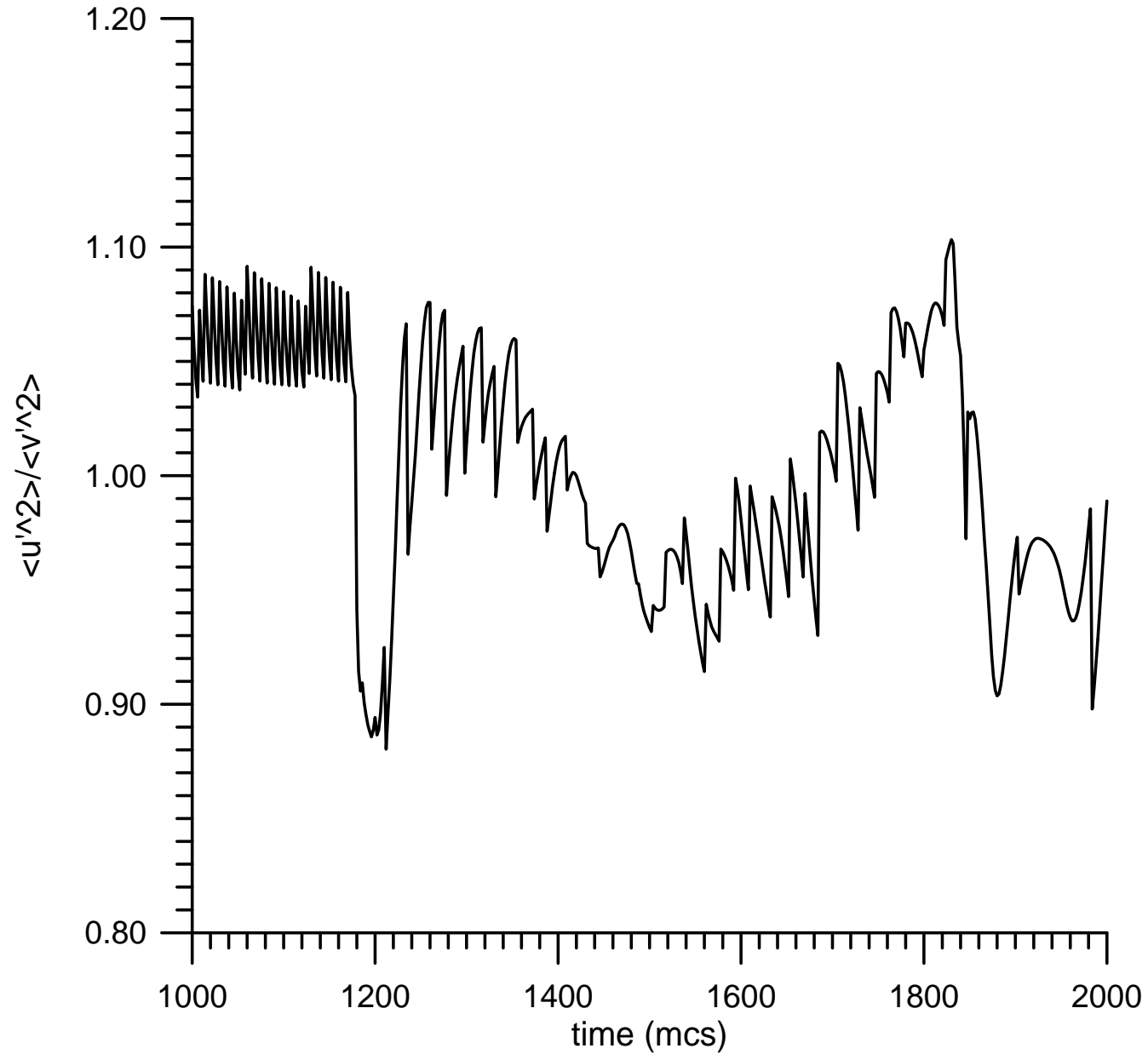
$\langle w'^2 \rangle$  - (curve 1),  $\langle v'^2 \rangle$  - (curve 2)  $\langle u'^2 \rangle$  - (curve 3)

t=1200mks

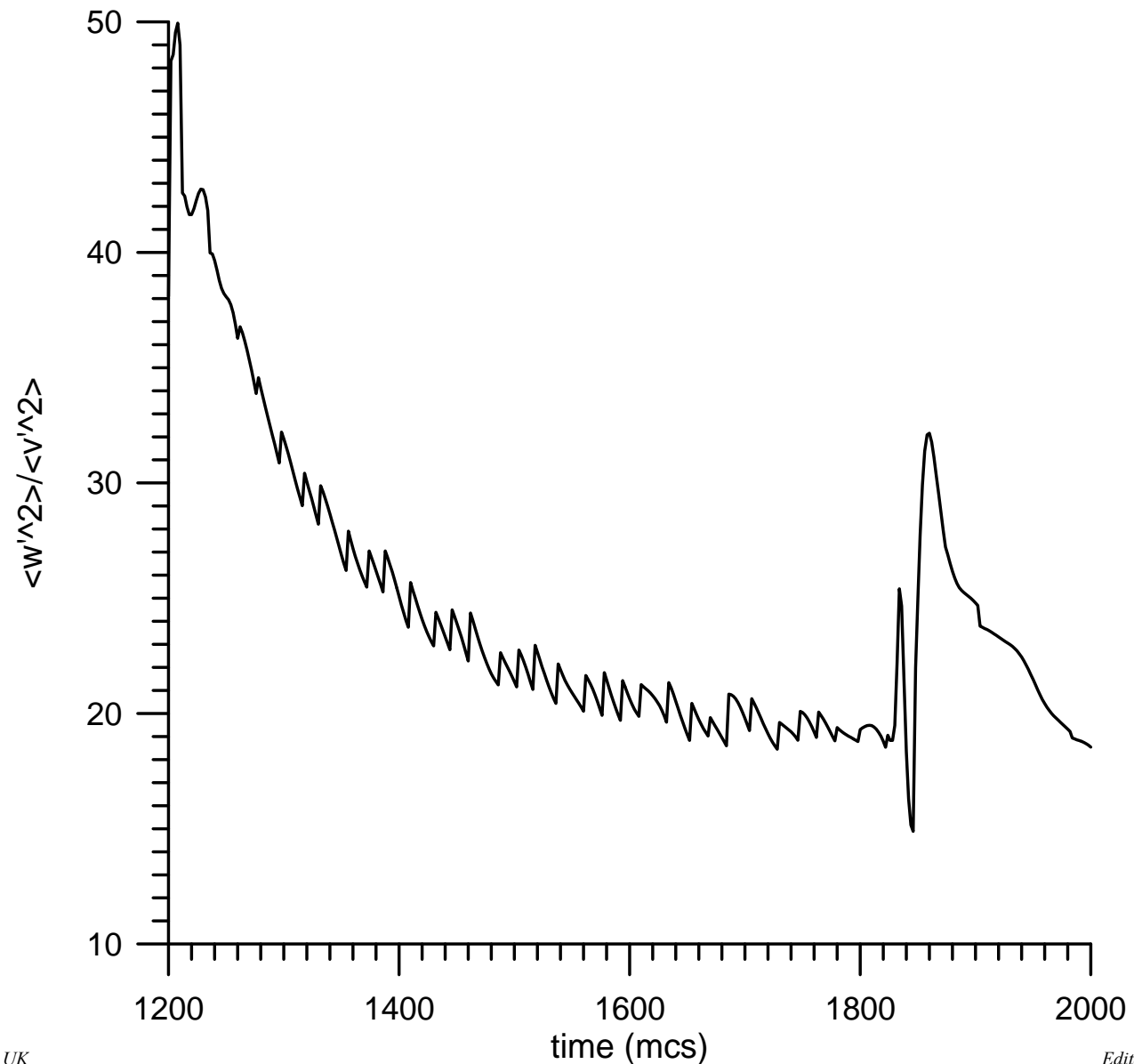
t=1600mks

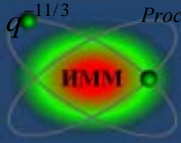


# Ratio of speed diametrical pulsation in task T2



$\langle v'^2 \rangle$  in task T2





# Conclusion

- Simulation of Richtmyer-Meshkov instability were fulfilled (parallel version of the NUT code) at a set of the condensed grids. ( $h=1.0, 0.5$  and  $1/3\text{mm}$ ).
- Dependence of mixing zone and velocity pulsations are obtain.
- The main characteristics of a mixing zone width growth with time and level of pulsation are established in a connection with the forms and amplitudes of the initial perturbations.
- The spectral analysis of the turbulent kinetic energy shows, that we observe inertial interval ( $E \sim q^{-5/3}$ ) and dissipative interval ( $E \sim q^{-11/3}$ )