

# **A New Two-Scale Mix Model: Towards a Multi-Component Model of Turbulent Mixing\***

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Physics of Compressible Turbulent Mixing**



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# Abstract



Turbulent mixing of the fluids in a multi-component system is of interest in situations such as inertial confinement fusion (ICF) and core-collapse supernovae<sup>1</sup>. We report results of a project to include a model of turbulent mixing in a multi-component hydrodynamics and physics model called KULL, which is used for ICF. Because KULL is a complex, multi-dimensional model, we have developed a simplified, one-dimensional version called sKULL to speed-up the development of the turbulent mixing model.

Of primary interest in the development of a turbulent mixing model for a multi-component fluid is the question of whether it is necessary to allow each component of the fluid to retain its own velocity. Generally a multi-component, multi-velocity turbulent mixing model should allow separate

velocities for each component of the fluid<sup>2</sup>. However, the necessity to carry separate velocities for each component of the fluid greatly increases the memory requirements and complexity of the computer implementation. In contrast, we present a new two-scale formulation of the K- $\epsilon$  turbulent mixing model, with production terms based on a recent scaling analysis<sup>3</sup>, which treats all components of the fluid as if they had the same velocity. We also show that our new method for the initial conditions of the uncoupled two-scale K- $\epsilon$  model yields asymptotic growth. Future work will compare the results of using this single velocity model with those from a more complete multi-velocity formulation of turbulent mixing, to decide whether the multi-velocity formulation needs to be used in KULL.

# The goal of this work is to develop a turbulent mixing model for the ICF code called KULL



- Turbulent mixing of the fluids in a multi-component system is of interest in situations such as inertial confinement fusion (ICF) and core-collapse supernovae<sup>1</sup>
- We report results of a project to include a model of turbulent mixing in a multi-component hydrodynamics and physics model called **KULL**, which is used for ICF
- Because **KULL** is a complex, multi-dimensional code, we have developed a simplified, one-dimensional version called **sKULL** to speed-up the development of the turbulent mixing model

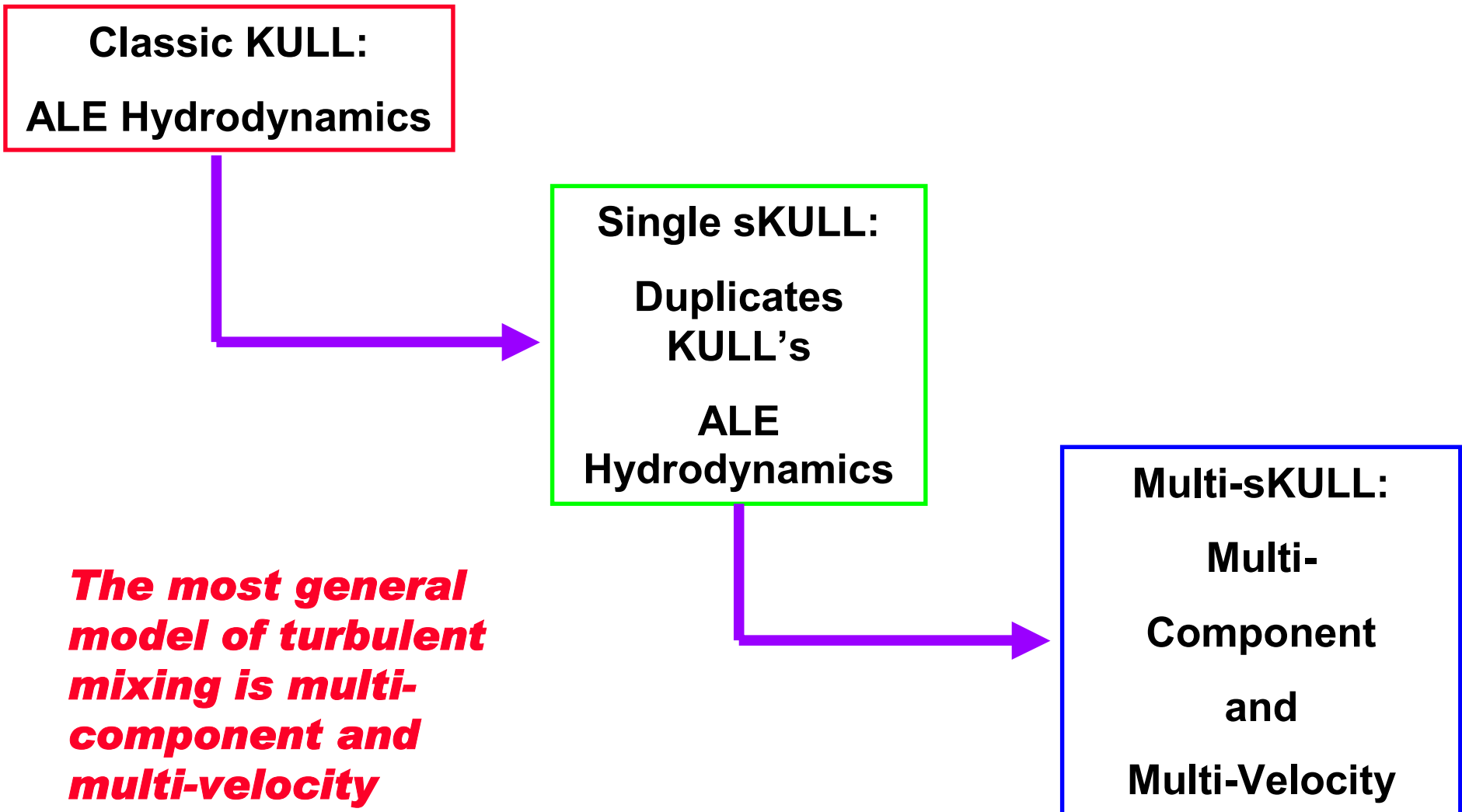
<sup>1</sup>Remington, B.A., D. Arnett, R.P. Drake, and H. Takabe, Modeling Astrophysical Phenomena in the Laboratory with Intense Lasers, *Science* **284**, 1488 (1999).

## Three areas of this research are highlighted



- **sKULL** reproduces KULL's multi-component hydrodynamics and numerics
- A single velocity, multi-component, **two-scale K- $\epsilon$**  turbulent mixing model has been developed within sKULL
- A new method for the uncoupled two-scale K- $\epsilon$  **initial conditions** yields asymptotic growth

# We have an appropriate path to develop a turbulent mixing model for KULL

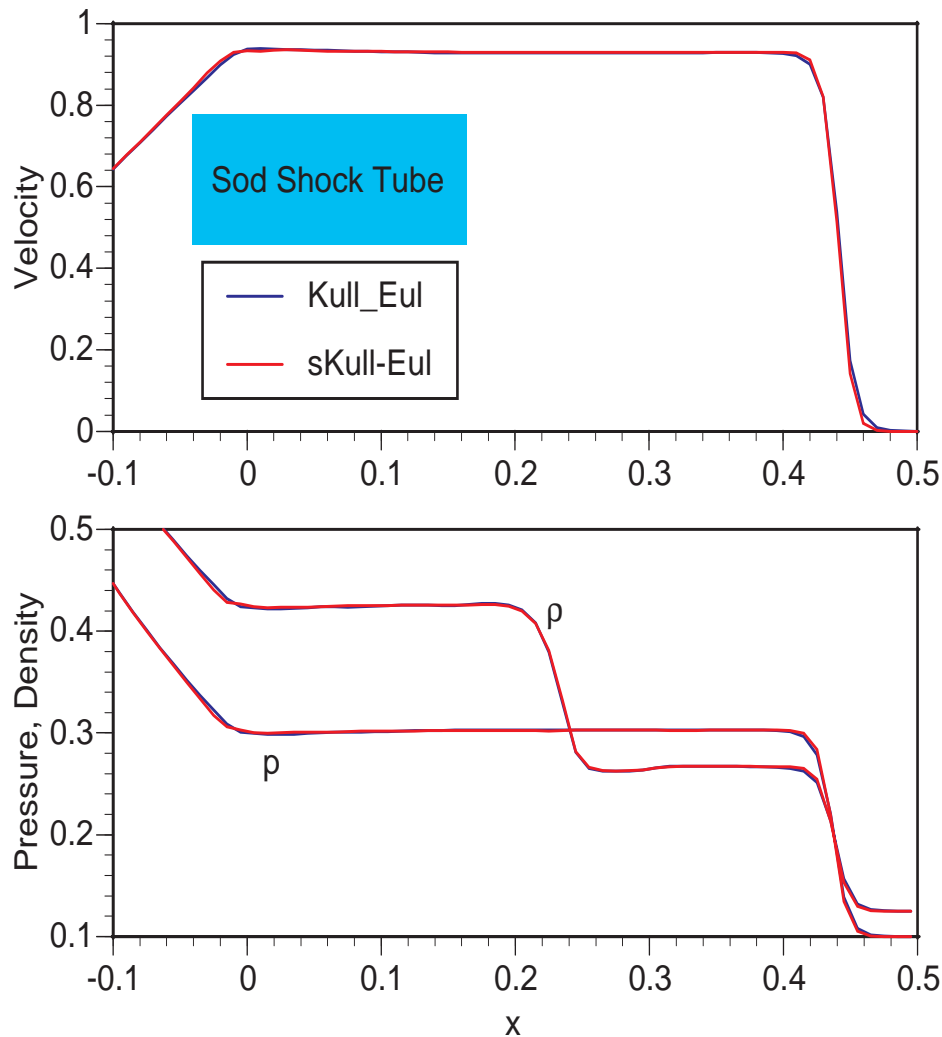


# **sKULL is the right platform in which to develop a turbulent mixing model for KULL**



- **sKULL duplicates KULL's hydrodynamics**
  - **Side-by-side runs of KULL and sKULL on the Sod shock produce the same results**
  - **We tested the Lagrangian, Eulerian, and ALE capabilities of sKULL to ensure they matched KULL's**
  
- **The simplified nature of sKULL, due both to 1-D and no additional physics, allows it to run more quickly**
  - **Faster run times lead to shorter turn-around times for testing turbulent mixing models**

# Side-by-side runs of KULL and sKULL on the Sod shock problem produce the same results



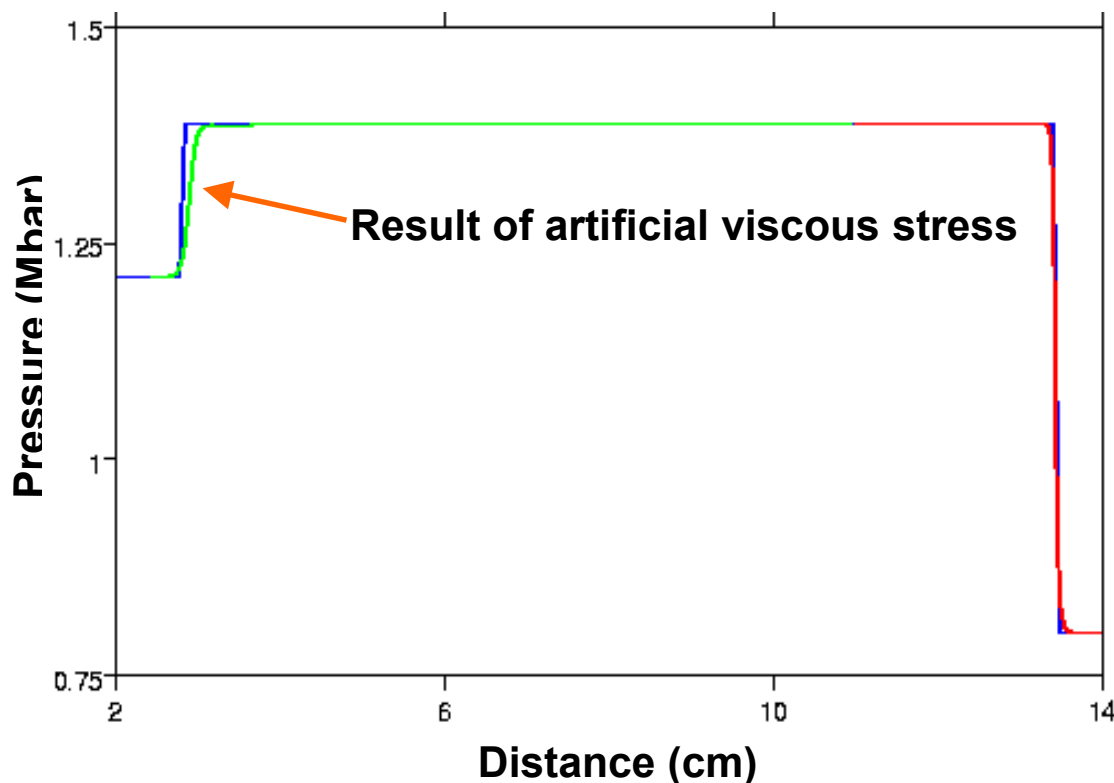
- Duplication of KULL results on selected problems w/ sKULL verifies that we've duplicated KULL's numerics

- Sod (1978) shock tube problem:

$\rho = 1$	$\rho = 0.125$
$p = 1$	$p = 0.1$
$u = 0$	$u = 0$

- Standard test problem
- Compared Lagrangian, Eulerian, and ALE results to ensure that the results from the two codes agreed

# sKULL MC-1V's simulation of the Benjamin air-SF<sub>6</sub> shock tube agrees well with the exact solution



- Benjamin *et al.* (1993) air-SF<sub>6</sub> shock tube:

Air*	SF <sub>6</sub>
$\rho = 1.27 \times 10^{-3}$	$\rho = 4.85 \times 10^{-3}$
$u = 1.05 \times 10^4$	$u = 0$
$p = 1.21 \times 10^6$	$p = 8.00 \times 10^5$

**Shock (Ma=1.2)**

- Pressure results from the MC-1V Lagrangian simulation versus exact solution at time 232  $\mu$ s show good agreement



## **A multi-component, multi-velocity (MC-MV) approach needs to be considered for the turbulent mixing model**



- In RTI/RMI, zones may contain more than one component, each with its own velocity
  - Component interactions (e.g., drag) can lead to mixing
  - From the rocket rig experiments, this led David Youngs (AWE) to create his MC-MV mixing model<sup>2</sup>
- The MC-MV equations add a great deal of complexity
  - Carrying separate velocities increases the memory requirement
  - The drag term may require an implicit treatment

<sup>2</sup>Youngs, D.L., *Laser & Particle Beams* 12, 725 (1994).

## **sKULL will be used to test multi-velocity versus single velocity-based turbulent mixing models**



**Because of sKULL's simplified nature it is faster and cheaper than KULL**

- **The extra memory requirement of MC-MV will be manageable**
- **Additional computation for interactions will be do-able**
- **Different numerical treatments of the drag term can be tested (explicit vs. implicit vs. iterated)**

***Faster and cheaper makes sKULL the ideal platform to test whether MC-MV might be needed in KULL***

# The MC-MV equations (Youngs<sup>2</sup>) add a great deal of complexity



ALE: “Grid Velocity”  $u = \begin{cases} u_r, \text{ Lagrangian} \\ 0, \text{ Eulerian} \end{cases}$

$$\frac{\partial x}{\partial t} = u$$

“ALE”-like,  
 $g = 0$  if shock tube,  
 Interactions,  
 Turbulence Transport

$$\frac{\partial (f_r \rho_r)}{\partial t} = - \frac{\partial}{\partial x} [f_r \rho_r (u_r - u)] - f_r \rho_r \frac{\partial u}{\partial x}$$

$$\frac{\partial (f_r \rho_r u_r)}{\partial t} = - \frac{\partial}{\partial x} [f_r \rho_r u_r (u_r - u)] - f_r \rho_r u_r \frac{\partial u}{\partial x} - f_r \frac{\partial P}{\partial x} + f_r \rho_r g$$

$$+ \sum_s (D_{rs} + M_{rs}) - m_r \frac{\partial \tau}{\partial x}$$

$$\frac{\partial (f_r \rho_r e_r)}{\partial t} = - \frac{\partial}{\partial x} [f_r \rho_r e_r (u_r - u)] - f_r \rho_r e_r \frac{\partial u}{\partial x} - h_r P_r \frac{\partial \bar{u}}{\partial x}$$

$$+ \frac{\partial}{\partial x} \left( f_r \rho_r v_r \frac{\partial e_r}{\partial x} \right) + f_r \varepsilon$$

Because of the MC-MV equations' complexity, we've first developed a single velocity version, MC-1V



sKULL MC-1V  
Lagrangian  
equations with a  
new two-scale K-ε  
mixing model

$$\frac{dx}{dt} = u$$

$$\frac{dV_r}{dt} \approx h_r \frac{dV}{dt}, V = \sum_r V_r$$

Reynolds  
stress

$$\rho \frac{Du}{Dt} = -\frac{\partial P}{\partial x} - \frac{\partial \tau}{\partial x}$$

Turbulent dissipation

$$f_r \rho_r \frac{De_r}{Dt} = -h_r P_r \frac{\partial u}{\partial x} + f_r \rho_r \epsilon_t + \frac{\partial}{\partial x} \left( f_r \rho_r \frac{v_r}{\sigma_e} \frac{\partial e_r}{\partial x} \right)$$

Kinematic viscosity/Schmidt number

# The use of the compressibility in the effective pressure allows the simplification to single velocity



$$P = \frac{\sum_r P_r f_r K_r}{\sum_r f_r K_r}, P_r = p_r + q_r$$

**Effective pressure  
(includes artificial  
viscosity)**

$$K_r^{-1} = \rho_r \left. \frac{\partial P_r}{\partial \rho_r} \right|_{e_r} + \frac{P_r}{\rho_r} \left. \frac{\partial P_r}{\partial e_r} \right|_{\rho_r}, h_r = \frac{f_r K_r}{\sum_s f_s K_s}$$

**Inverse effective  
compression;  
Relative com-  
pression**

**For an ideal gas and  $q_r = 0$ ,  $K_r^{-1} = \gamma_r p_r$  (adiabatic compressibility), and  $h_r = [f_r / (\rho_r c_r^2)] / [\sum_s f_s / (\rho_s c_s^2)]$  (Youngs<sup>2</sup>)**

# The viscosity for energy diffusion and Reynolds stress comes from the two-scale K-ε model



$$\tau = \frac{2}{3} \rho K_p - \frac{4}{3} \rho \nu_p \frac{\partial u}{\partial x} \quad \text{Reynolds stress}$$

$$\nu_\alpha = \nu_0 + \nu_{T\alpha} \quad \text{Kinematic visc.}$$

Molecular

Turbulent

$$\nu_{T\alpha} = C_\mu \frac{K_\alpha^2}{\varepsilon_\alpha}$$

Equations for  $K_\alpha$  and  $\varepsilon_\alpha$  are needed for closure

# The two-scale K-ε equations describe evolution of the production and turbulence scales



$$\frac{DK_p}{Dt} = P_{R^*} - \varepsilon_p + \frac{\partial v_p}{\partial x} \frac{\partial K_p}{\sigma_K \partial x} - \frac{\tau}{\rho} \frac{\partial u}{\partial x} \quad \text{Production scale}$$

$$\frac{DK_t}{Dt} = \varepsilon_p - \varepsilon_t + \frac{\partial v_t}{\partial x} \frac{\partial K_t}{\sigma_K \partial x} \quad \text{Turbulence scale}$$

$$\frac{D\varepsilon_p}{Dt} = C_{p1} \frac{\varepsilon_p}{K_p} P_{R^*} - C_{p2} \frac{\varepsilon_p^2}{K_p} + \frac{\partial v_p}{\partial x} \frac{\partial \varepsilon_p}{\sigma_\varepsilon \partial x}$$

$$\frac{D\varepsilon_t}{Dt} = C_{t1} \frac{\varepsilon_p \varepsilon_t}{K_t} - C_{t2} \frac{\varepsilon_t^2}{K_t} + \frac{\partial v_t}{\partial x} \frac{\partial \varepsilon_t}{\sigma_\varepsilon \partial x}$$

# The production terms $P_{R^*}$ for the two-scale K- $\epsilon$ equations parameterize mixing caused by RTI or RMI



Based on a recent scaling analysis<sup>3</sup> of RT and RM instabilities, the production term may be written as

## Rayleigh-Taylor

$$P_{RT} = 4C_{RT}\epsilon_p^{1/2} (gA)^{3/4} (k_0^{-1/4} - k_1^{-1/4})$$

## Richtmyer-Meshkov

$$P_{RM} = 2C_{RM}\epsilon_p^{1/2} (A\Delta u)^{3/2} (k_1^{1/2} - k_0^{1/2})$$

<sup>3</sup>Zhou, Y., A scaling analysis of turbulent flows driven by Rayleigh-Taylor and Richtmyer-Meshkov instabilities, *Phys. Fluids* **13**, 538–543 (2001).



# Wave numbers $k_0$ and $k_1$ for the production terms evolve with the flow



Initially  $k_0$  and  $k_1$  are set by the initial perturbation scales, but thereafter evolve according to the computed production and turbulence scales

## Production scale

### Rayleigh-Taylor

$$k_0 = \left[ \frac{\frac{4}{7} C_{RT} \varepsilon_p^{1/2} (gA)^{1/4}}{\frac{4}{7} C_{RT} \varepsilon_p^{1/2} (gA)^{1/4} k_1^{-3/4} + K_p} \right]^{4/3}$$

### Richtmyer-Meshkov

$$k_0 = \frac{4 C_{RM}^2 \varepsilon_p A \Delta u}{\left[ 2 C_{RM} (\varepsilon_p A \Delta u)^{1/2} k_1^{-1/2} + K_p \right]^2}$$

## Turbulence scale

$$k_1 = \varepsilon_t \left( \frac{3}{2} C_K / K_t \right)^{3/2} \quad \text{RT or RM}$$

The change in total energy is due to production minus dissipation and surface fluxes



$$\underbrace{\frac{D}{Dt} \int d(u^2/2 + e + K)}_{\text{Total Energy Change}} = \underbrace{\int d(P_{R^*} - \varepsilon_t)}_{\text{Production - Dissipation}} - \underbrace{[u(P + \tau) + F_e + F_K]}_{\text{Surface Fluxes}} \int dS$$

**Total Energy Change**

**Production - Dissipation**

**Surface Fluxes**

$$F_e = -\frac{\nu}{\sigma_K} \frac{\partial e}{\partial x}, F_{K_\alpha} = -\frac{\nu_\alpha}{\sigma_K} \frac{\partial K_\alpha}{\partial x}$$

**Diffusive fluxes of internal and turbulent kinetic energies**

$$F_K = F_{K_p} + F_{K_t}$$

# Results from Orszag and Speziale will be used to provide ICs for the two-scale turbulent mixing model



- **Steve Orszag's work for the ASCI Turbulence Group:**

$$K_p < \tilde{P}_{R0} \sqrt{\nu_0 / C_\mu}, \tilde{P}_{R0} = \boxed{P_{R0}} / \mathcal{E}_{p0}^{1/2}$$

**Initial RT or RM Production**

$$\frac{C_\mu K_{p0}^2}{\nu_0} < \mathcal{E}_{p0} < \tilde{P}_{R0}^2$$

**If violated: too much turbulence initially, interface dies out**

**If violated: no turbulent viscosity develops, Orszag's high Re run blew up**

- **Is this result consistent w/ Speziale's fixed point analysis?**

## The result using Orszag's approach is consistent with Speziale's fixed point analysis



- The “fixed points” from Speziale's analysis<sup>4</sup> act as attractors
- Initialize with fixed points that are consistent with desired long-term behavior
  - ↓ Leads more quickly to the desired long-term state
- Speziale's analysis yields the following fixed points:

$$\mathcal{E}_{pf} = f_p P_{R0}, \mathcal{E}_{tf} = f_t \mathcal{E}_{pf}$$

$$f_p = (C_{p1} - 1) / (C_{p2} - 1), f_t = (C_{t1} - 1) / (C_{t2} - 1)$$

suggesting

$$\mathcal{E}_{p0} = f_p^2 \tilde{P}_{R0}^2$$

↓ consistent with Orszag's approach if  $f_p < 1$

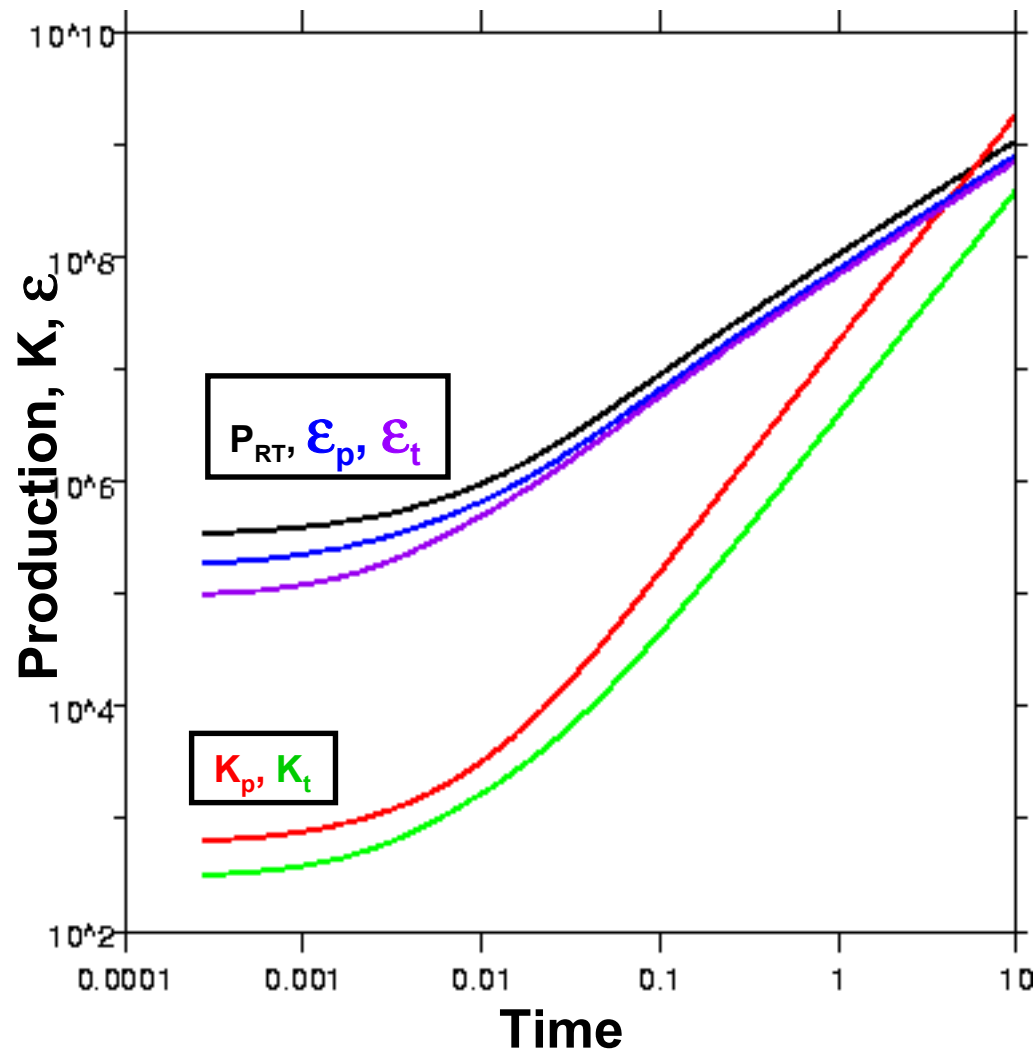
<sup>4</sup>Speziale, C.G., and N. Mac Giolla Mhuiris, On the prediction of equilibrium states in homogeneous turbulence, *J. Fluid Mech.*, **209**, 591-615 (1989).

## The Orszag-Speziale ICs yield asymptotically growing solutions for the two-scale turbulent mixing model



- **Current recommended values  $C_{p1} = 1.5$ ,  $C_{p2} = 2$  give  $f_p = 1/2$** 
  - ↓ **The result using Orszag's approach is consistent with Speziale's fixed point analysis**
  - ↓ **Also consistent with constraints  $C_{*2} > 1$  and  $C_{*2} > C_{*1} > 1/2$  needed to get asymptotically growing solutions**
- **Solution of the two-scale turbulent mixing model equations as ODEs using the Orszag-Speziale ICs produced the expected asymptotically growing results**
  - ↓ **The Orszag-Speziale ICs will be used in the two-scale turbulent mixing model for sKULL MC-1V**

# A test of the Orszag-Speziale ICs w/ the uncoupled two-scale K- $\epsilon$ model yields asymptotic growth



## Rayleigh-Taylor Test

- $gA = 4000$
- $k_1(0) = 10k_0(0) = 20\pi$
- $C_K = C_{RT} = 1.5$
- $C_{p1} = 1.5, C_{p2} = 2.0$
- $C_{t1} = 1.08, C_{t2} = 1.15$
- Constraints for asymptotic growth are satisfied, and asymptotic growth achieved

Consistency with a stand-alone ODE solution demonstrates that the uncoupled two-scale K- $\epsilon$  equations have been correctly implemented in sKULL

## Summary and Future Work



- A simplified version of the ICF code KULL has been developed which reproduces KULL's multi-component hydrodynamics
  - The purpose of simplified KULL, sKULL, is to serve as a test-bed for implementation of multi-component turbulent mixing models
  - Tests show that sKULL faithfully duplicates KULL's numerics
- A single velocity, multi-component, two-scale K- $\epsilon$  turbulent mixing model has been developed within sKULL
  - A new method for the uncoupled two-scale K- $\epsilon$  initial conditions yields asymptotic growth
- Future work will compare these single velocity results with those from a more complete multi-velocity formulation of turbulent mixing, to decide whether the multi-velocity formulation needs to be used in KULL

## Table of Symbols #1



<u>Quantity</u>	<u>Description</u>
$D_{rs}$	Drag force on fluid r due to fluid s
$e_r$	Specific internal energy of fluid r
$\mathcal{E}_p$	Dissipation at production scale
$\mathcal{E}_t$	Dissipation at turbulence scale
$f_r$	Volume fraction of fluid r
$g$	Acceleration (e.g., gravitational)
$h_r$	Relative compressibility of fluid r
$K_p$	Turbulence kinetic energy at production scale
$K_t$	Turbulence kinetic energy at turbulence scale
$m_r$	Mass fraction of fluid r



## Table of Symbols #2



<b>Quantity</b>	<b>Description</b>
$M_{rs}$	Added mass effect for fluid r due to fluid s
$\nu_{0r}$	Molecular viscosity of fluid r
$\nu_T = \nu_p + \nu_t$	Total turbulent viscosity
$\nu_p$	Turbulent viscosity, production scale
$\nu_t$	Turbulent viscosity, turbulence scale
$\nu_r = \nu_{0r} + \nu_T$	Total viscosity of fluid r
$P_r = p_r + q_r$	Effective pressure of fluid r
$P = p + q$	Effective mean pressure
$p = \sum_r h_r p_r$	Mean pressure
$p_r = p_r(\rho_r, e_r)$	Pressure of fluid r from EOS

## Table of Symbols #3



<u>Quantity</u>	<u>Description</u>
$q = \sum_r f_r q_r$	Mean artificial viscous stress
$q_r$	Artificial viscous stress of fluid r
$\rho = \sum_r f_r \rho_r$	Mean density
$\rho_r$	Density of fluid r
$t$	Time
$u$	Mean velocity
$\bar{u} = \sum_r f_r \bar{u}_r$	Vdume-weighted mean velocity
$\bar{u}_r = u_r + \frac{v_r}{\rho_r} \frac{\partial \rho_r}{\partial x}$	Vdume-weighted velocity of fluid r
$u_r$	Velocity of fluid r
$x$	Position at time t